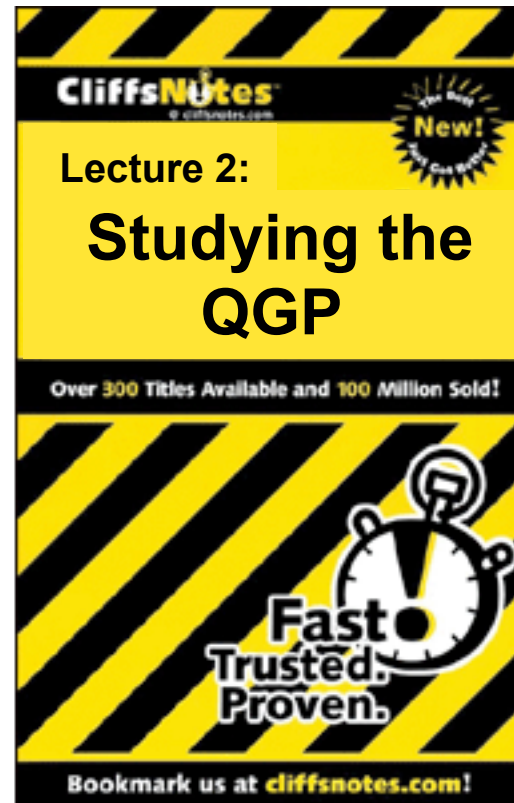
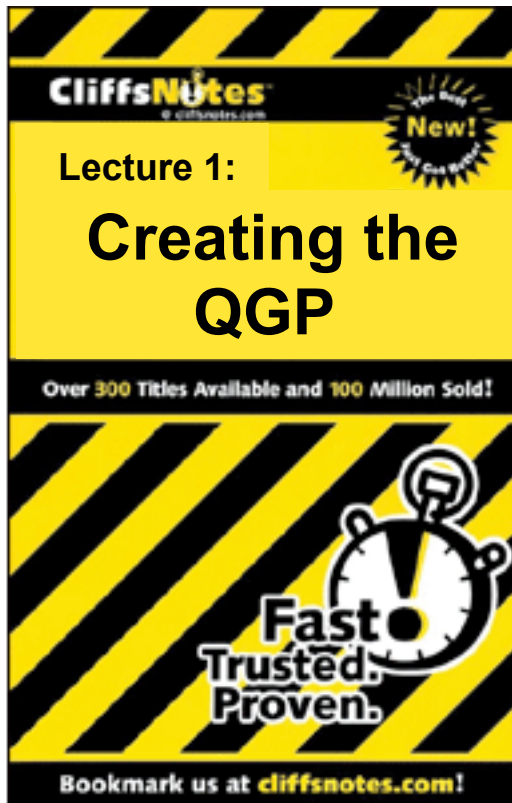


Studying a new Phase of Matter - An introduction to the Quark Gluon Plasma and Relativistic Heavy Ion Physics



2010 Hadron Collider Physics Summer School

Fermi Lab - August 2010

Helen Caines - Yale University

Relativistic Heavy Ions I - The What, Where, Why, and How of It All

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Summer School*

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Outline :

QCD and Asymptotic Freedom

The Quark Gluon Plasma

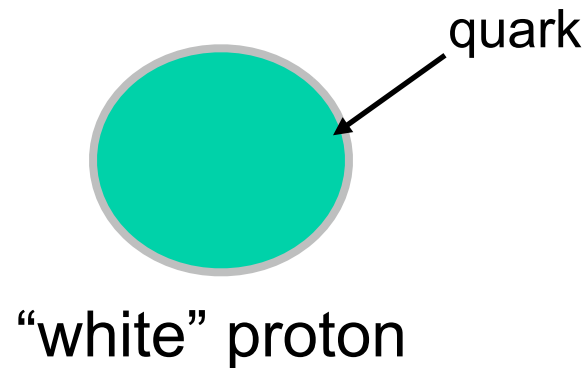
The Accelerators & Experiments

Evidence for the QGP



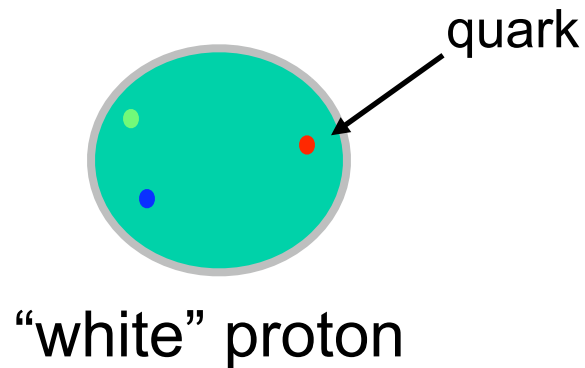
Confinement - QCD

Confinement: fundamental & crucial feature of strong interaction
force = const has significant consequences



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Confinement - QCD

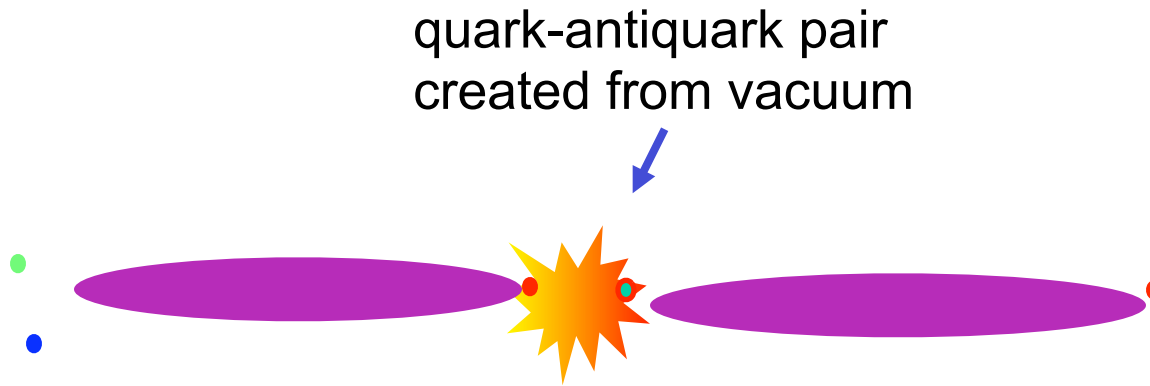
Confinement: fundamental & crucial feature of strong interaction
force = const has significant consequences



Strong **color** field
Force *grows* with
separation !!!

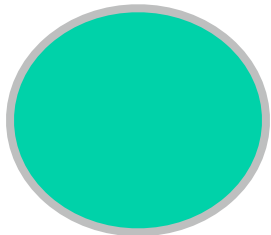
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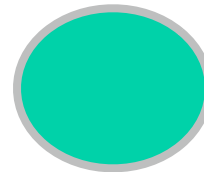


Confinement - QCD

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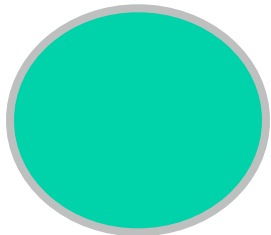
“white” proton
(confined quarks)



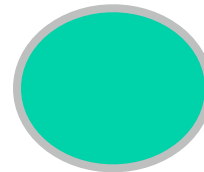
“white” π^0
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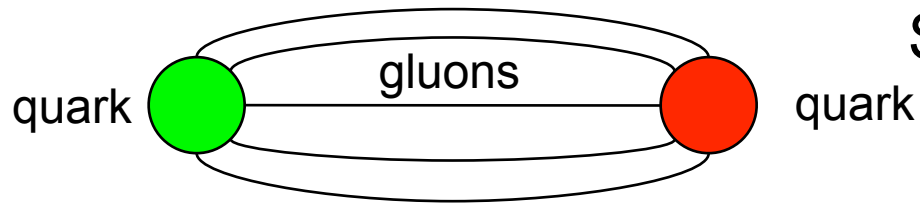
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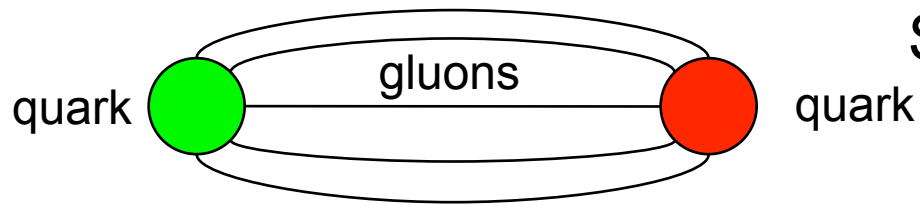
To understand the strong force and confinement: Create and study a system of deconfined colored quarks and gluons

We don't see free quarks



Strong force becomes a constant at ~size of a hadron which is $\sim 1 \text{ fm}$ (10^{-15} m)

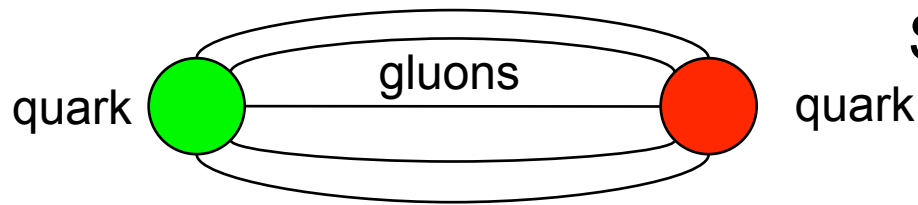
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Strong force becomes a constant at ~size of a hadron which is ~1 fm (10^{-15} m)

$$1 \frac{\text{GeV}}{\text{fm}} = \frac{10^9 \text{eV}}{10^{-15} \text{m}} \times \frac{1.6 \times 10^{-19} \text{J}}{\text{eV}} = 1.6 \times 10^5 \text{N}$$

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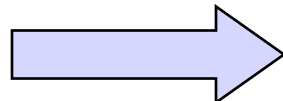


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Compare to gravitational force at Earth's surface

$$F = 1.6 \times 10^5 \text{N} = M \times g = M \times 9.8 \text{m/s}^2$$

 $M = 16,300 \text{kg}$

Quarks exert 16 metric tons of force on each other!

Asymptotic freedom

Coupling constant is not a “constant”

Runs with Q^2 (mtm transfer)
accounts for vacuum polarisation

$$\alpha_s(Q^2) = \frac{\alpha_s(\mu^2)}{[1 + (\alpha_s(\mu^2) \frac{(33-2n_f)}{12\pi}) \ln(Q^2/\mu^2)]}$$

$\alpha_s(\mu^2) \sim 1$!!

μ^2 : renormalization scale

33: gluon contribution

n_f : # quark flavours

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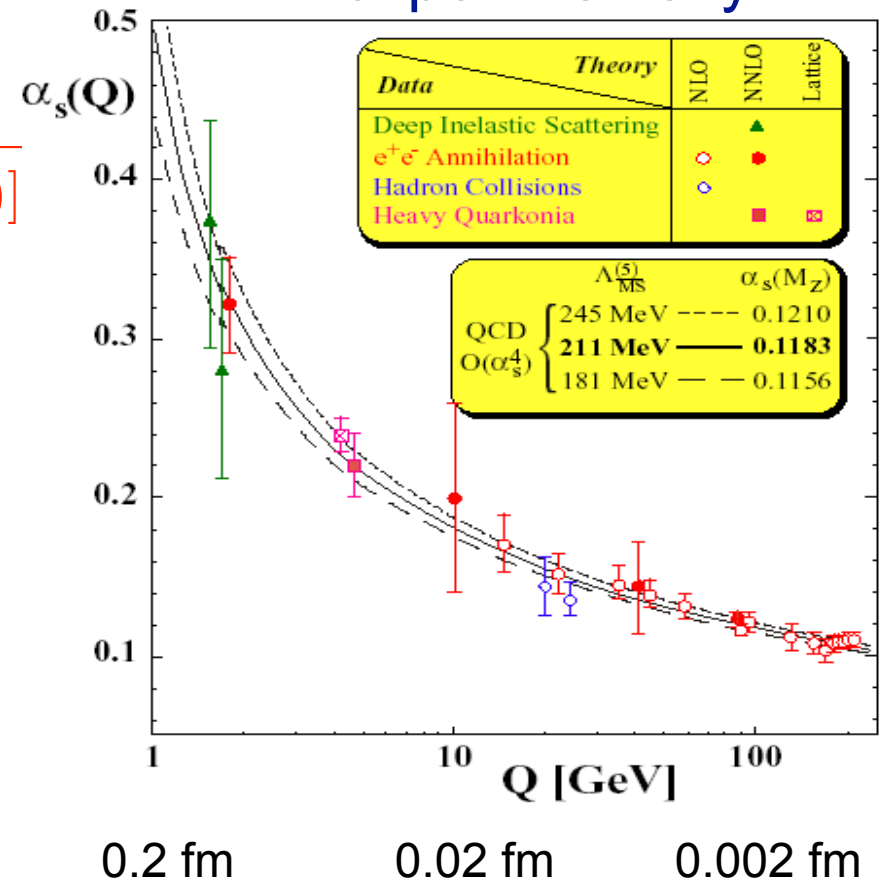
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Running measured
experimentally



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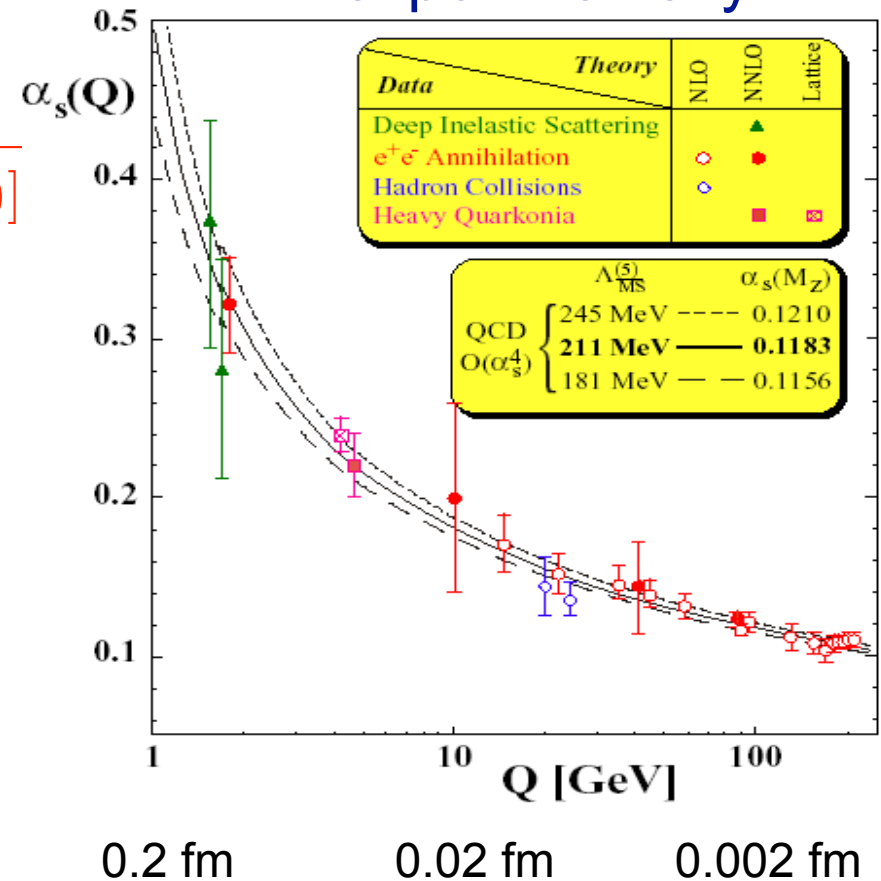
$\alpha_s(Q^2) \rightarrow 0$, as $Q \rightarrow \infty$, $r \rightarrow 0$

Coupling very weak

→ partons are essentially free

Asymptotic Freedom

Running measured
experimentally



Asymptotic freedom

Coupling constant is not a “constant”



The Nobel Prize in Physics 2004

“for the discovery of asymptotic freedom in the theory of the strong interaction”



David J. Gross



H. David Politzer



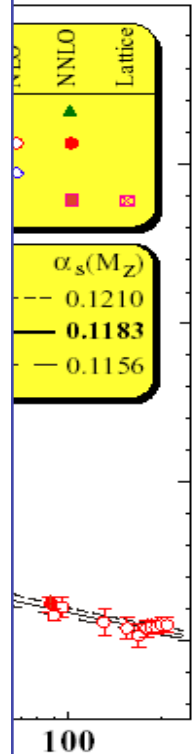
Frank Wilczek

Asymptotic Freedom

0.2 fm

0.02 fm

0.002 fm



Asymptotic freedom vs Debye screening

Asymptotic freedom occurs at very high Q^2

Problem: Q^2 much higher than available in the lab.

So how to create and study this new phase of matter?

Solution: Use effects of **Debye screening**

In the presence of many **colour** charges (charge density n), the **short** range term of the strong potential is modified:

Charges at long range ($r > r_D$) are screened

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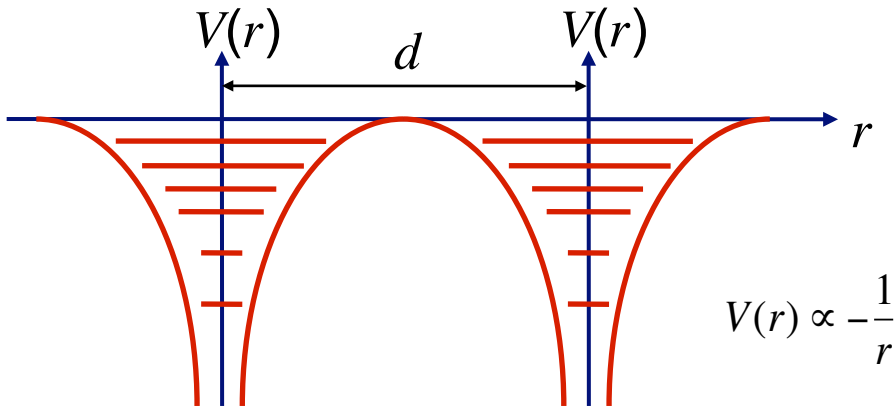
$$V_s(r) \propto \frac{1}{r} \implies \frac{1}{r} \exp\left[\frac{-r}{r_D}\right]$$

where $r_D = \frac{1}{3\sqrt{n}}$ is the **Debye radius**

Charges at long range ($r > r_D$) are screened

QED and Debye screening

$r > r_D$



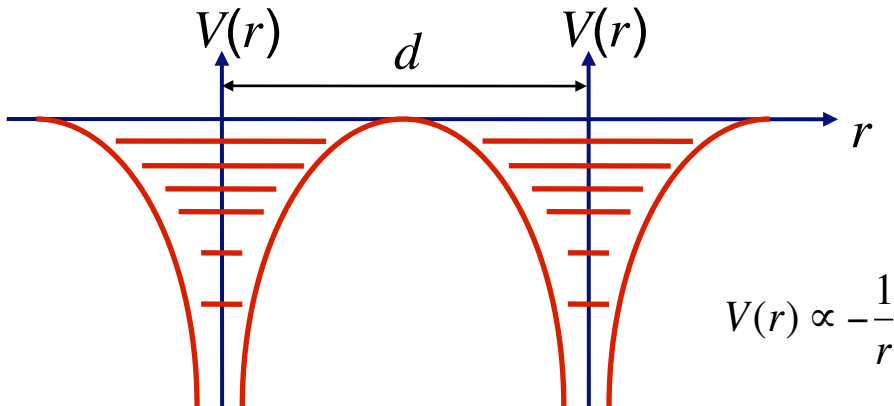
$r < r_D$

In condensed matter this leads to an interesting transition

e^- separation $>$ e^- binding radius
 \rightarrow insulator

QED and Debye screening

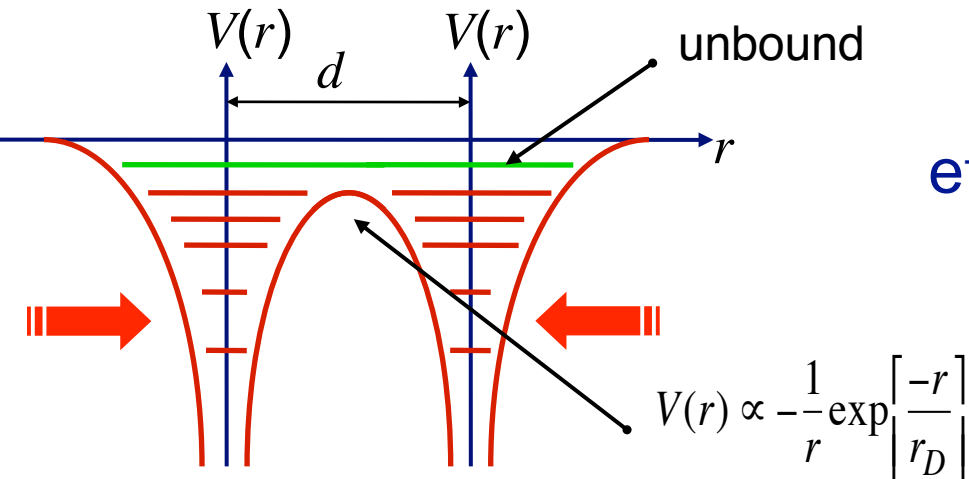
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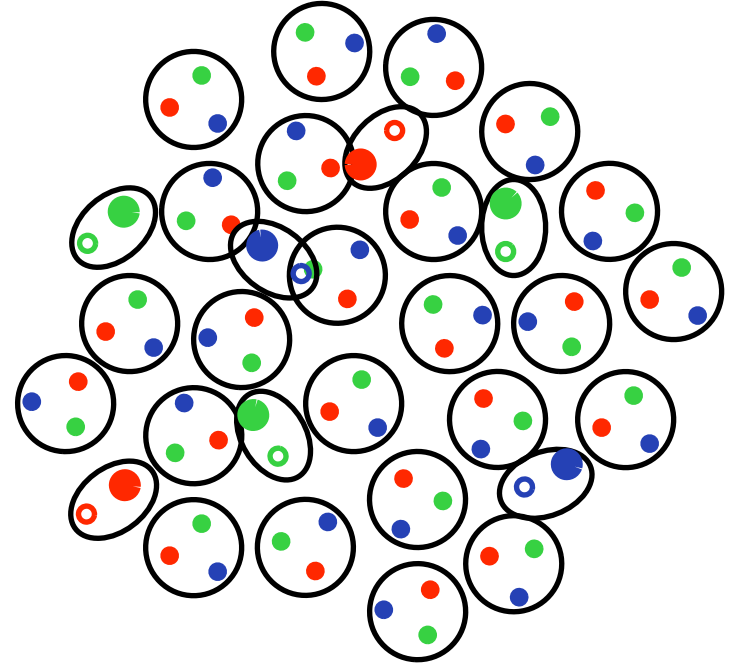
e^- separation $<$ e^- binding radius
 \rightarrow conductor

This is the Mott Transition

QCD and Debye screening

At low colour densities:

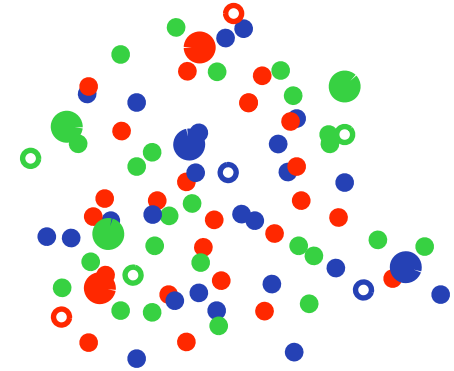
quarks and gluons confined into
colour singlets
→ hadrons (baryons and mesons)



QCD and Debye screening

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At high colour densities:

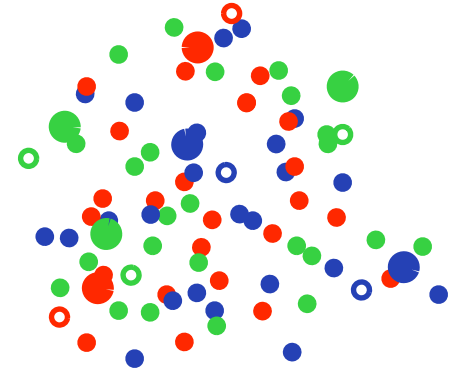
quarks and gluons unbound
Debye screening of colour charge

→ QGP - colour conductor

QCD and Debye screening

At low colour densities:

quarks and gluons confined into
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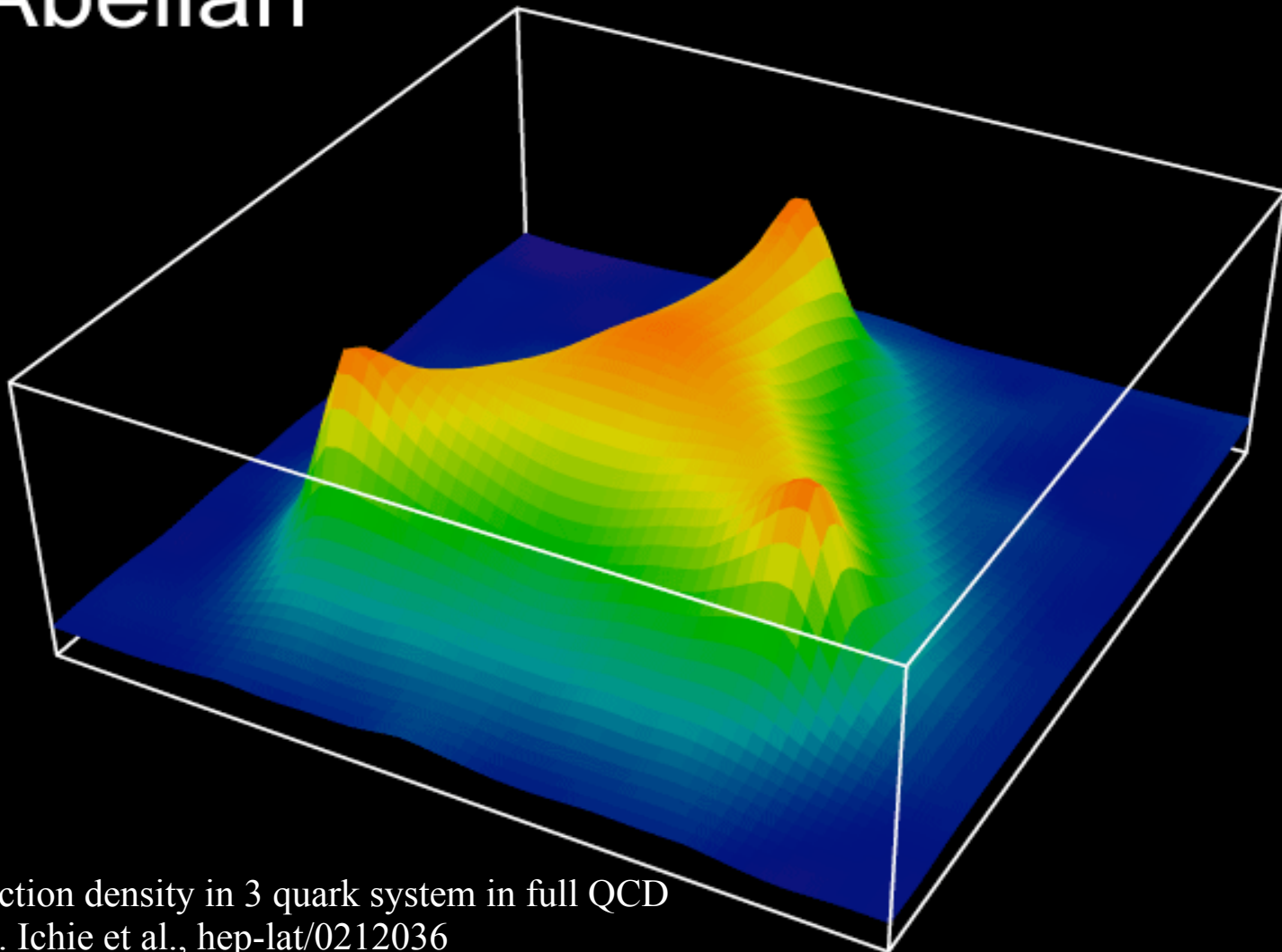
→ QGP - colour conductor

Can create high colour density by heating or compressing

→ QGP creation via accelerators or in neutron stars

What is T_c ? - Lattice QCD

Abelian

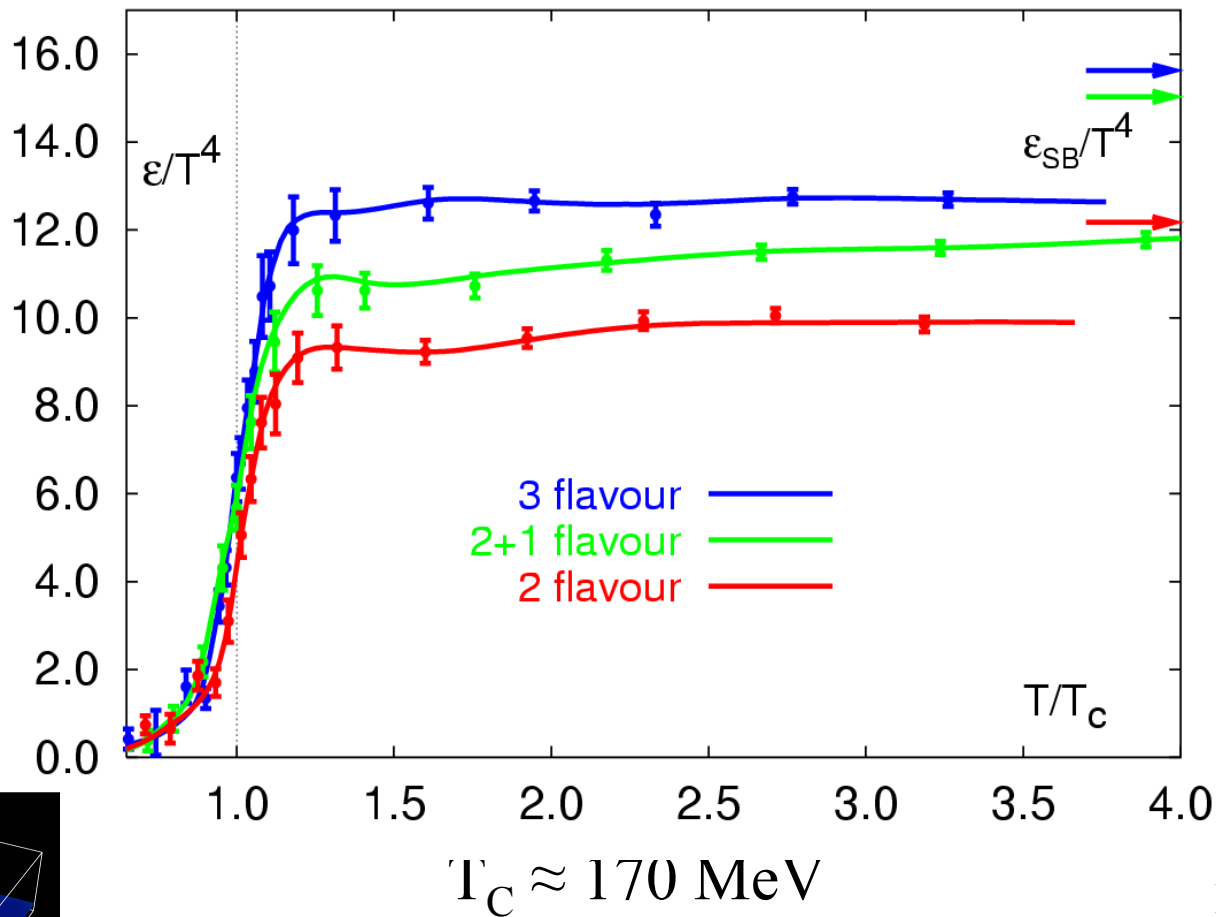


Action density in 3 quark system in full QCD
H. Ichie et al., hep-lat/0212036

Helen Caines - HPCSS - August 2010

What is T_c ? - Lattice QCD

- Coincident transitions: deconfinement and chiral symmetry restoration
- Recently extended to $\mu_B > 0$, order still unclear (1st, 2nd, crossover ?)



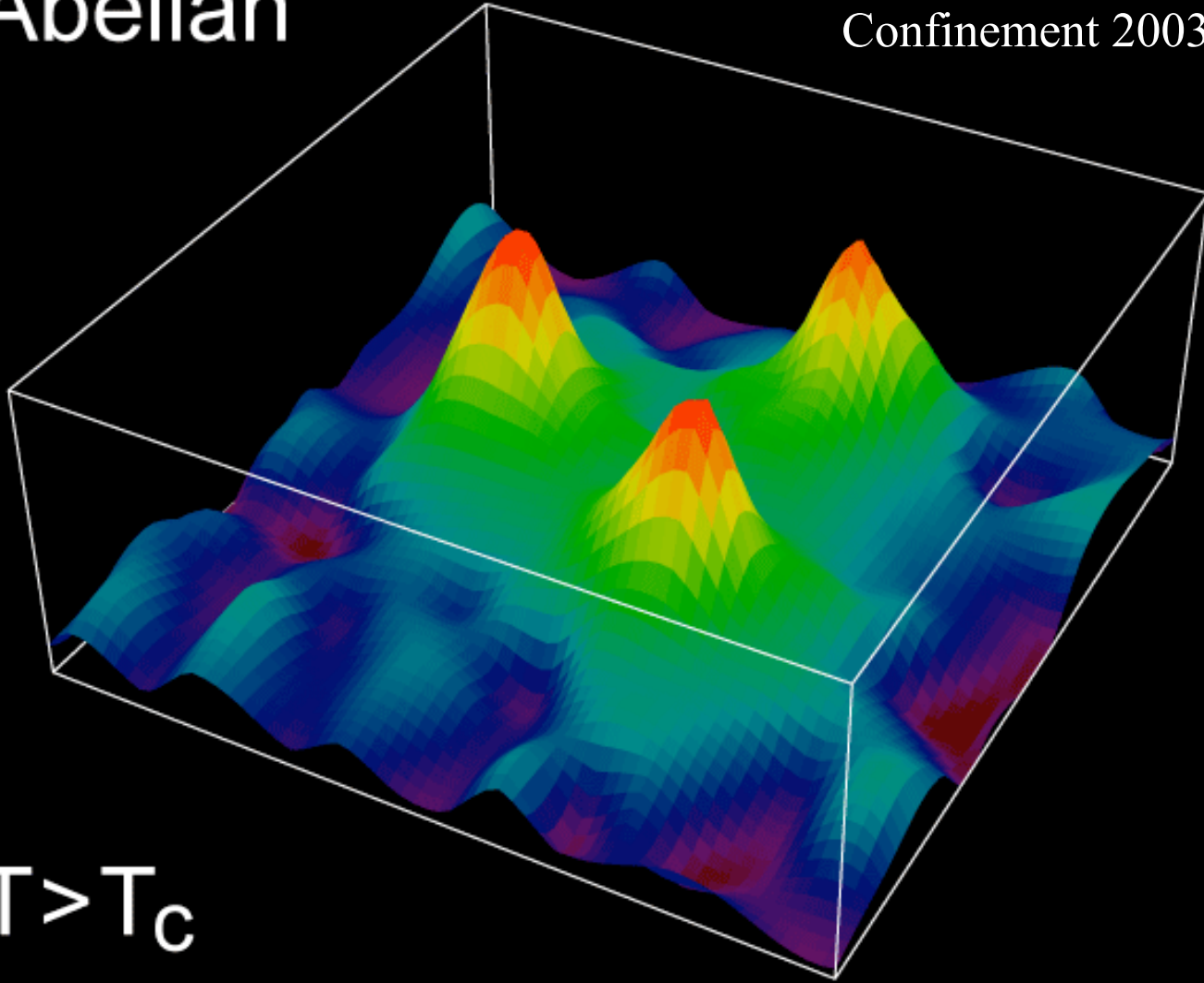
F. Karsch,
hep-ph/0103314

Helen Caines - HPCSS - August 2010

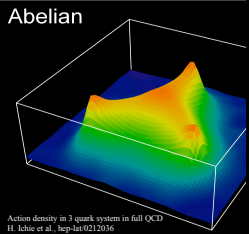
What is T_c ? - Lattice QCD

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G. Schierholz *et al.*,
Confinement 2003

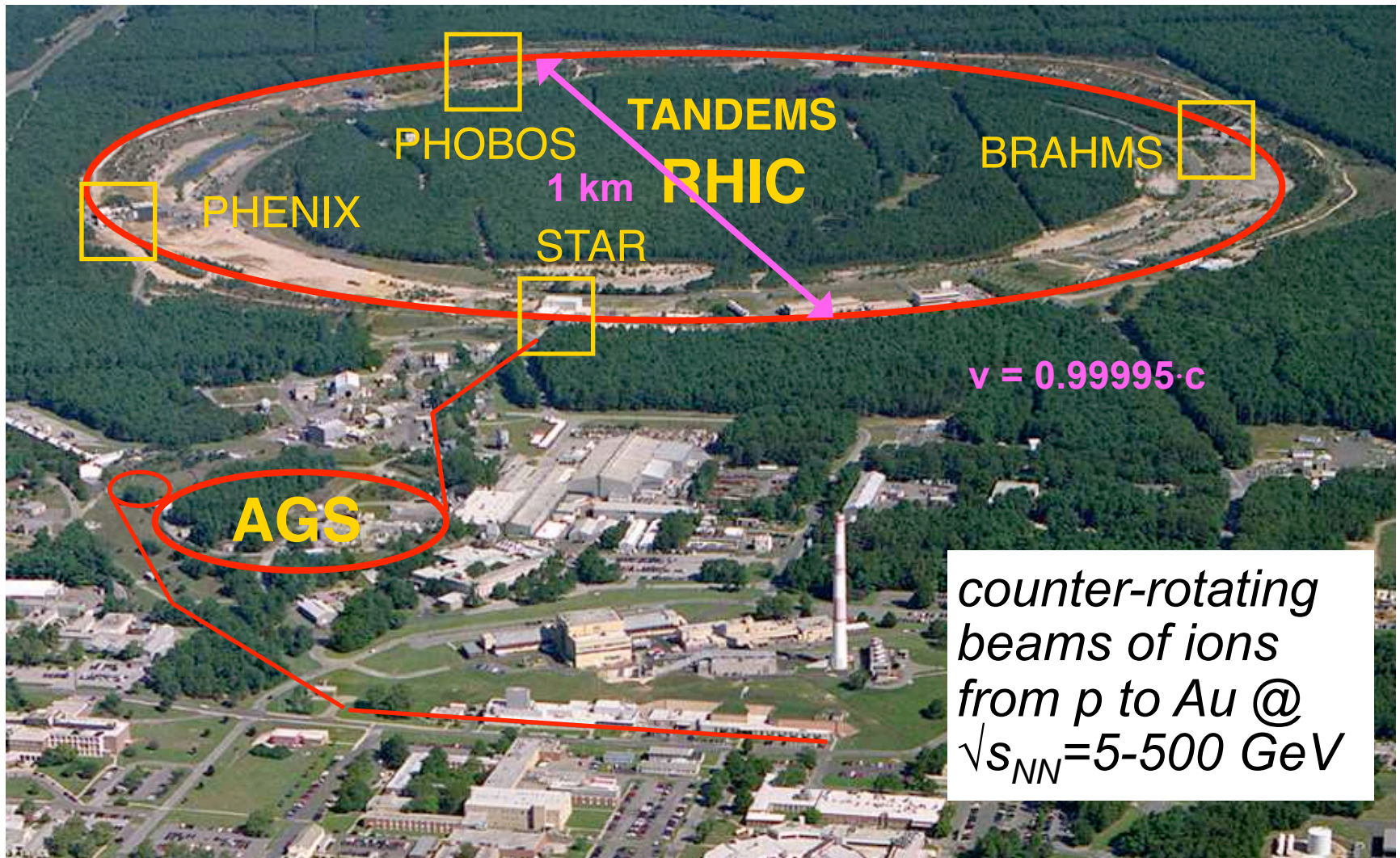


$T > T_c$

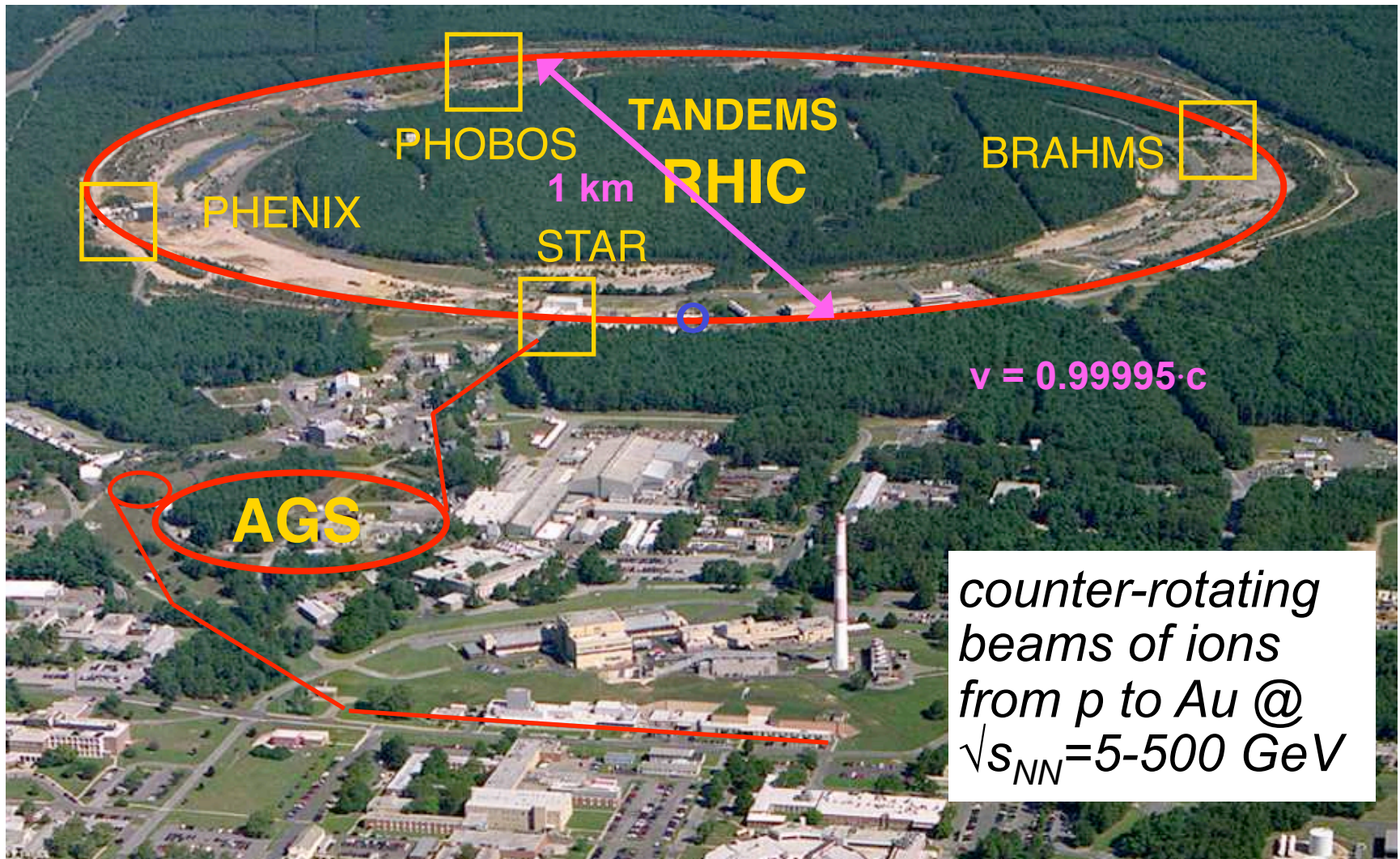


Helen Caines - HPCSS - August 2010

RHIC - a collider



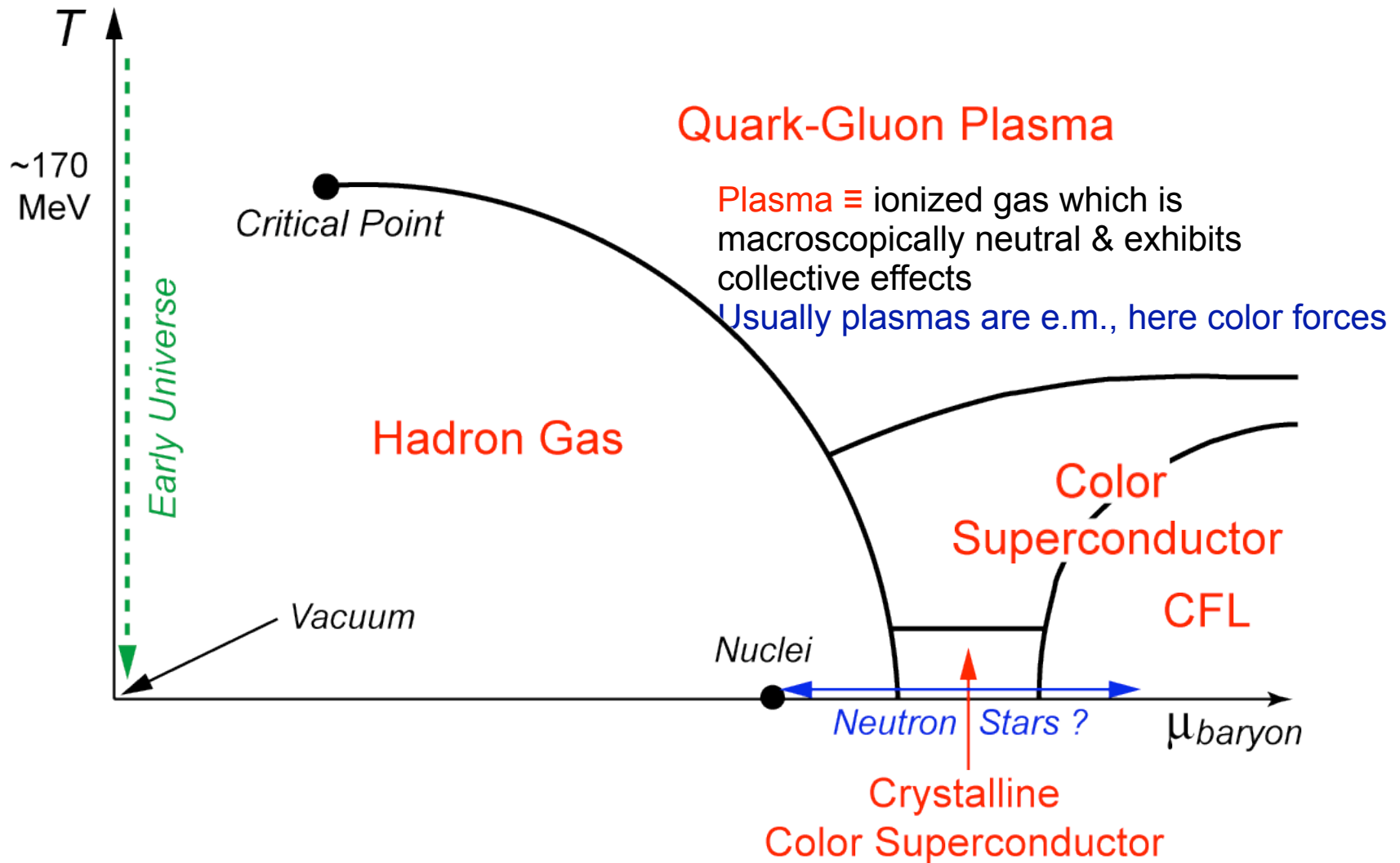
RHIC - a collider



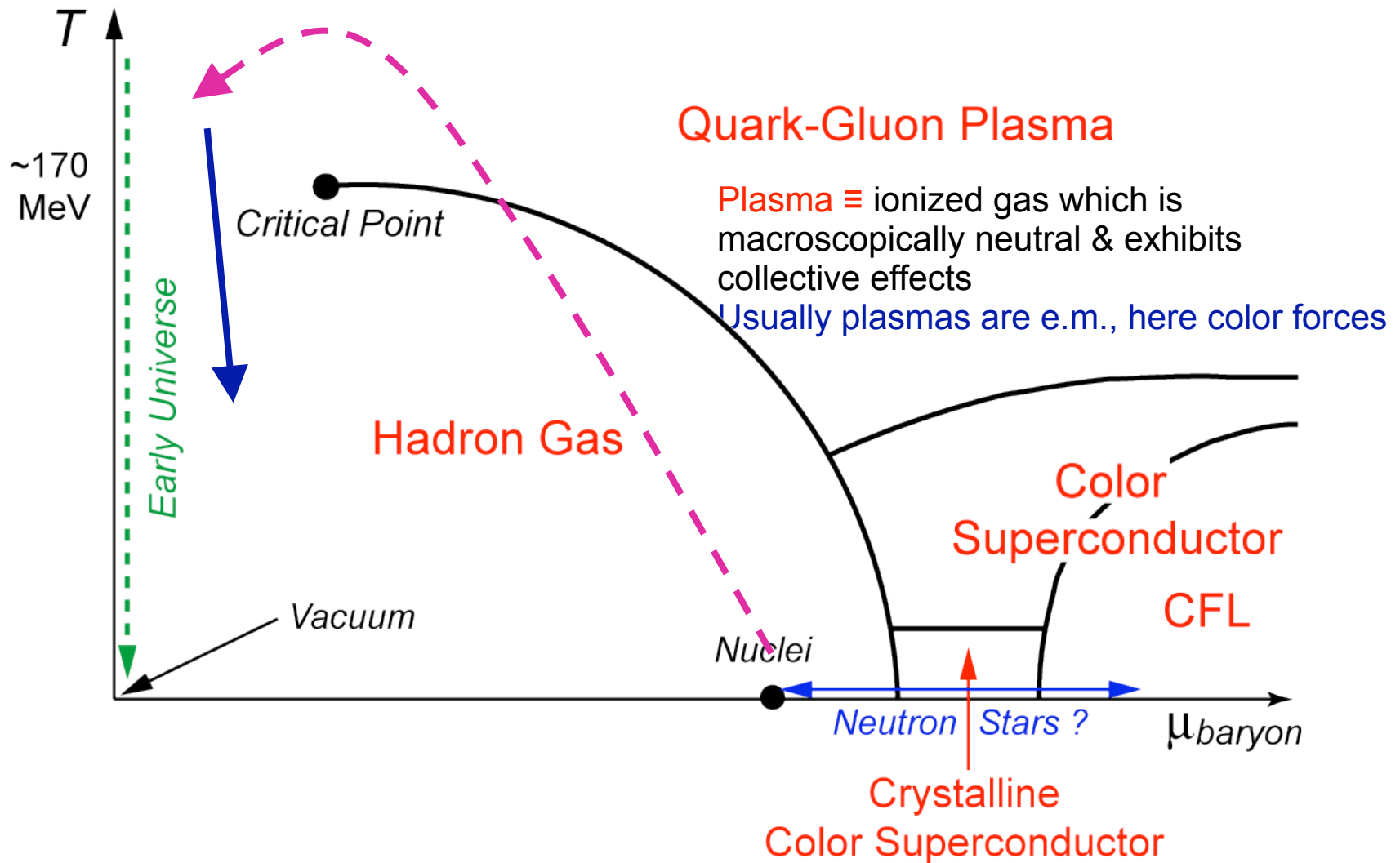
RHIC and the LHC

	RHIC	LHC
Start date	2001	2009
Ion	Au-Au & p-p	Pb-Pb & p-p
Max \sqrt{s}	200 GeV	5.5 TeV
Circumference	2.4 miles	17 miles
Depth	On surface	175 m below ground
HI Exp.	BRAHMS, PHENIX, PHOBOS, STAR	ALICE, ATLAS, CMS
Located	BNL, New York, USA	CERN, Geneva, Switzerland

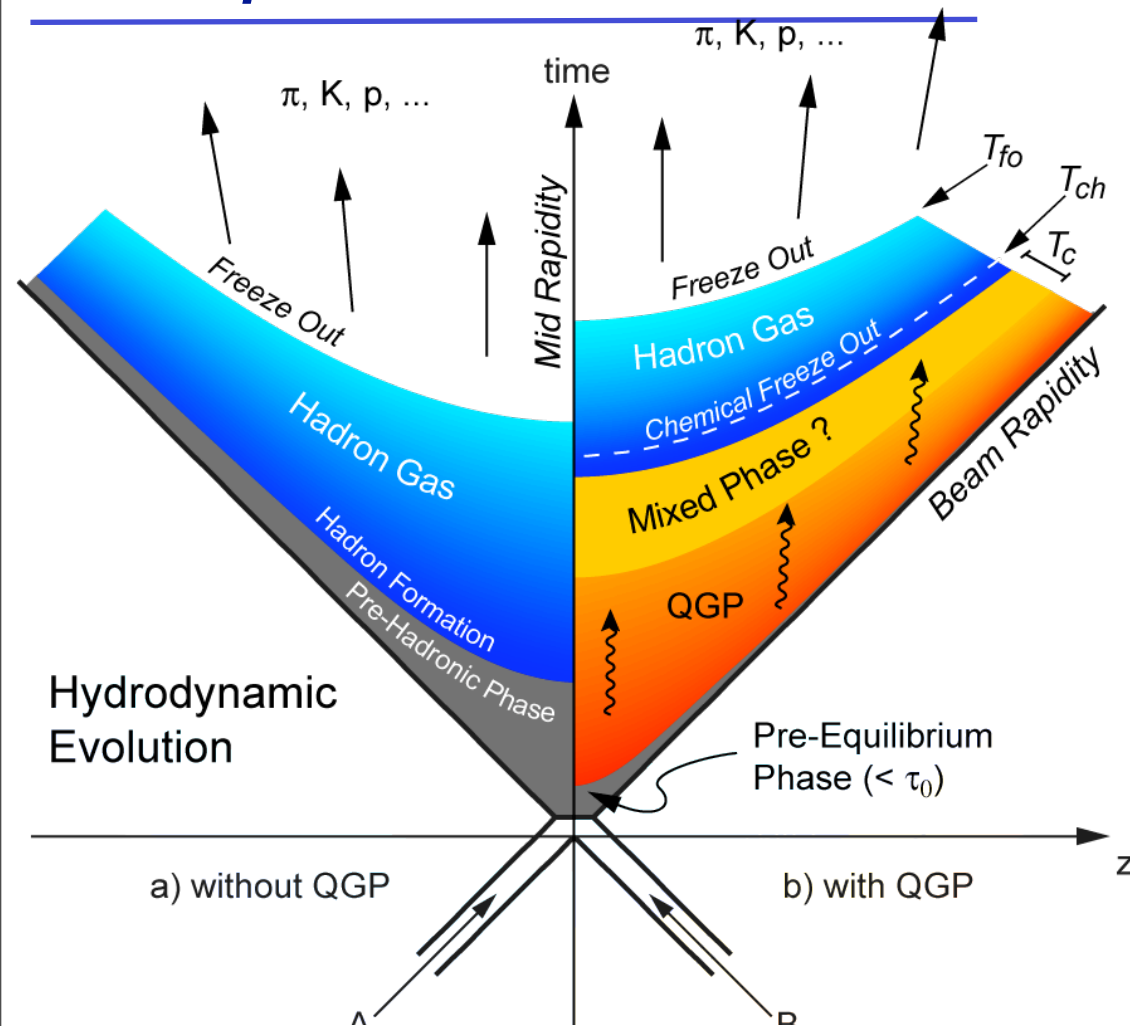
QCD phase diagram of hadronic matter



QCD phase diagram of hadronic matter



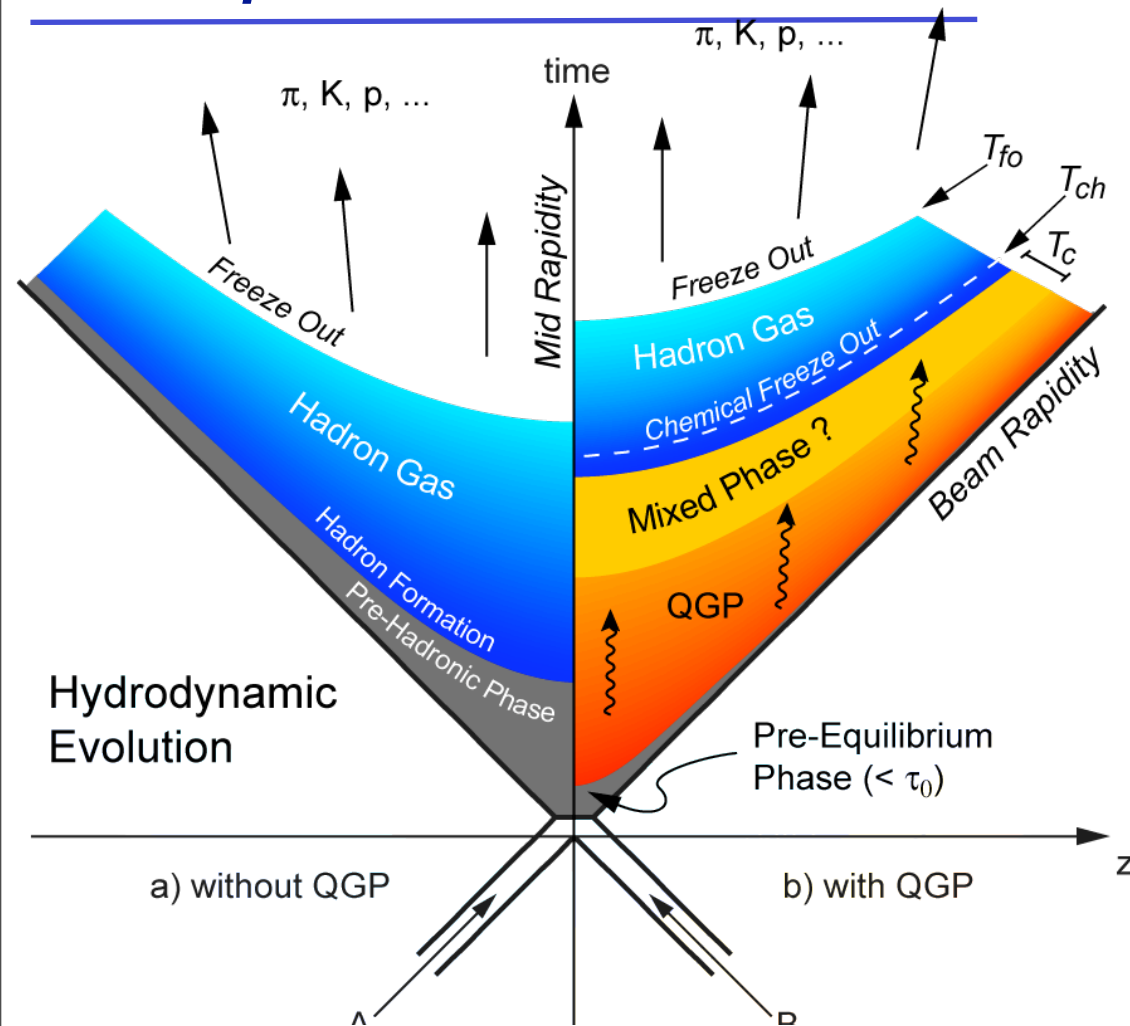
The phase transition in the laboratory



Chemical freeze-out ($T_{ch} \leq T_c$): inelastic scattering ceases

Kinetic freeze-out ($T_{fo} \leq T_{ch}$): elastic scattering ceases

The phase transition in the laboratory



Lattice (2-flavor):

$$T_C \approx 173 \pm 8 \text{ MeV}$$

$$\epsilon_C \approx (6 \pm 2) T^4 \approx 0.70 \text{ GeV/fm}^3$$

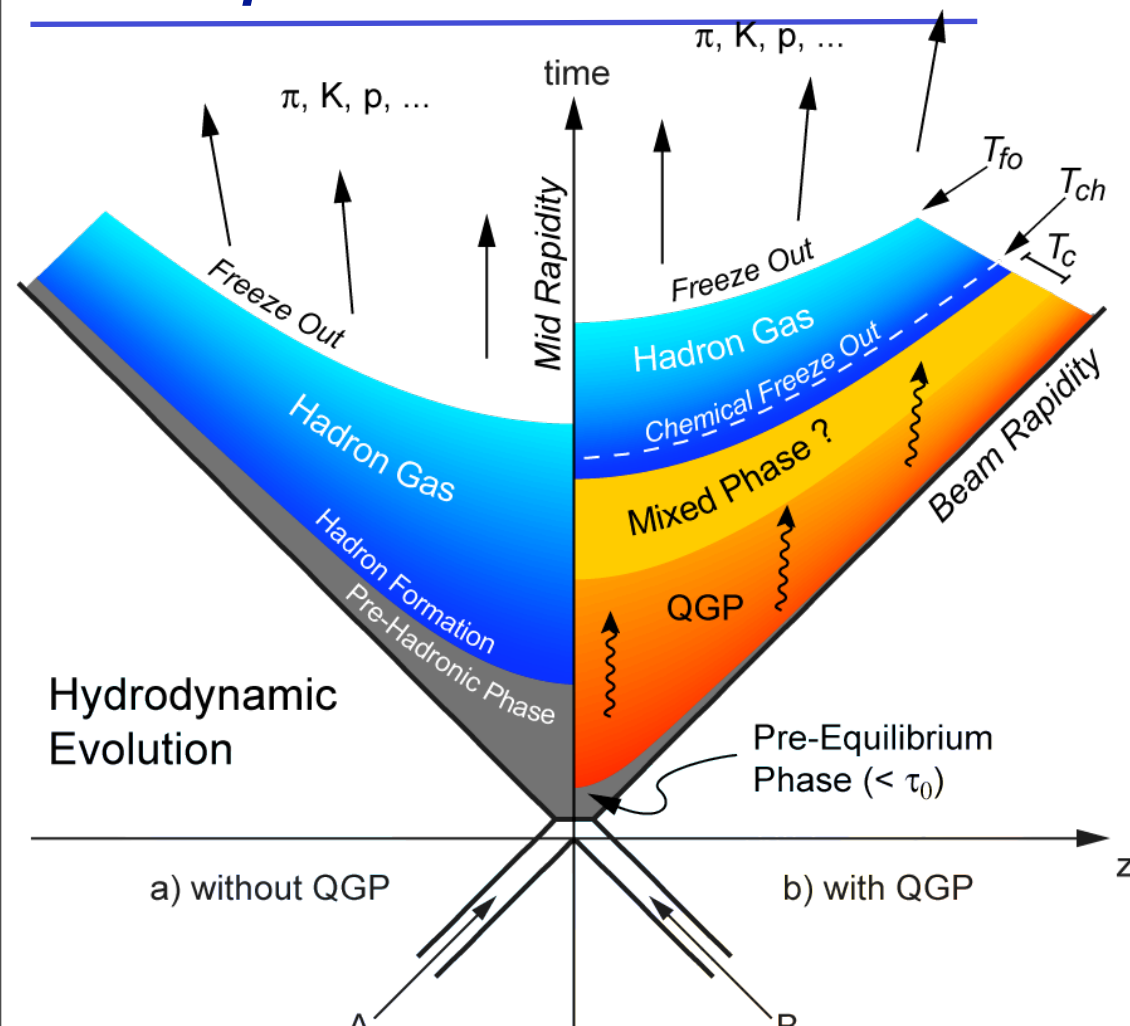
Remember: cold nuclear matter

$$\epsilon_{\text{cold}} \approx u / \frac{4}{3} \pi r_0^3 \approx 0.13 \text{ GeV/fm}^3$$

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Necessary but not sufficient condition

Tevatron (Fermilab)

$$\epsilon(\sqrt{s} = 1.8 \text{ TeV } pp) \gg$$

$$\epsilon(\sqrt{s} = 200 \text{ GeV Au+Au RHIC})$$

Thermal Equilibrium \Rightarrow many constituents

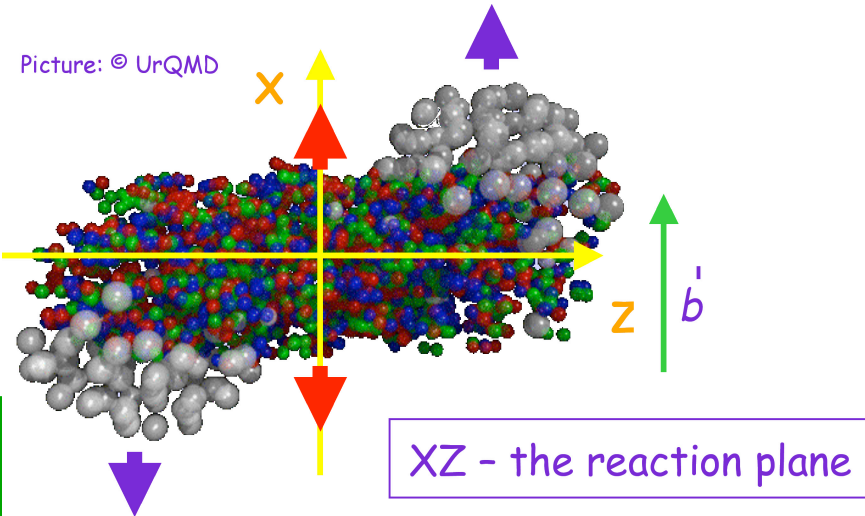
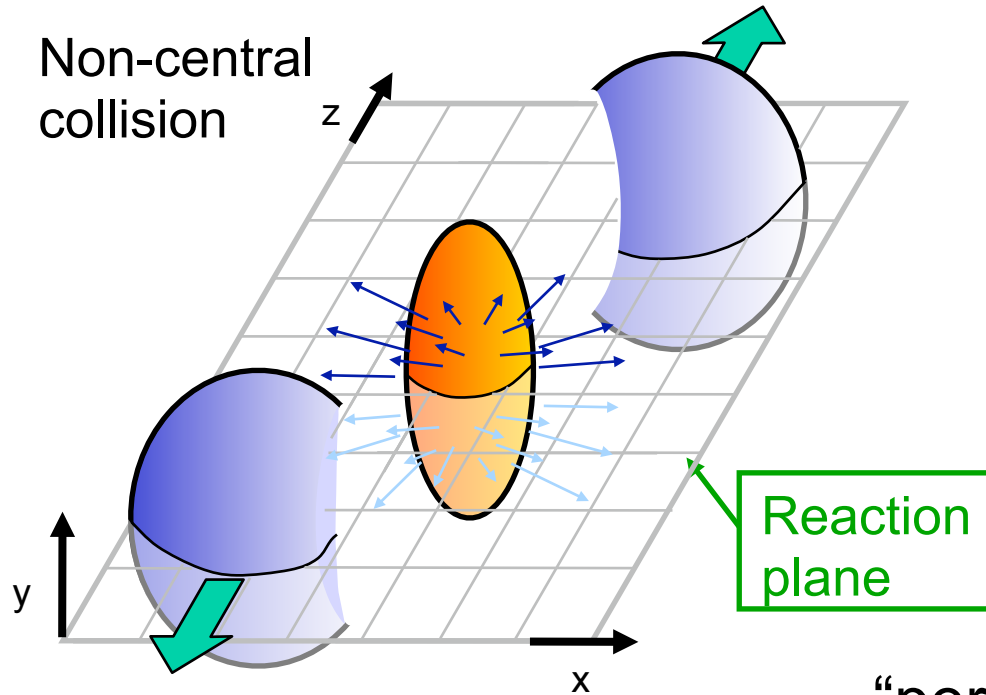
Chemical freeze-out ($T_{\text{ch}} \leq T_c$): inelastic scattering ceases

Kinetic freeze-out ($T_{\text{fo}} \leq T_{\text{ch}}$): elastic scattering ceases

Size matters !!!

Geometry of a heavy-ion collision

Non-central collision

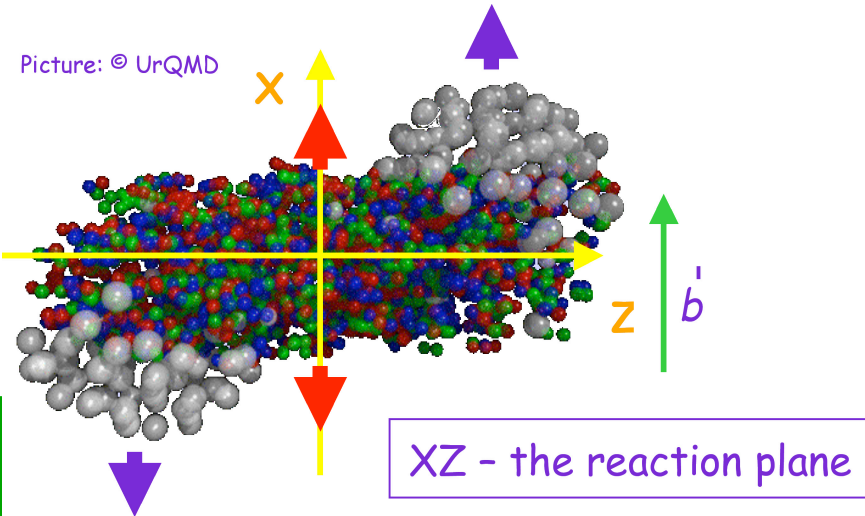
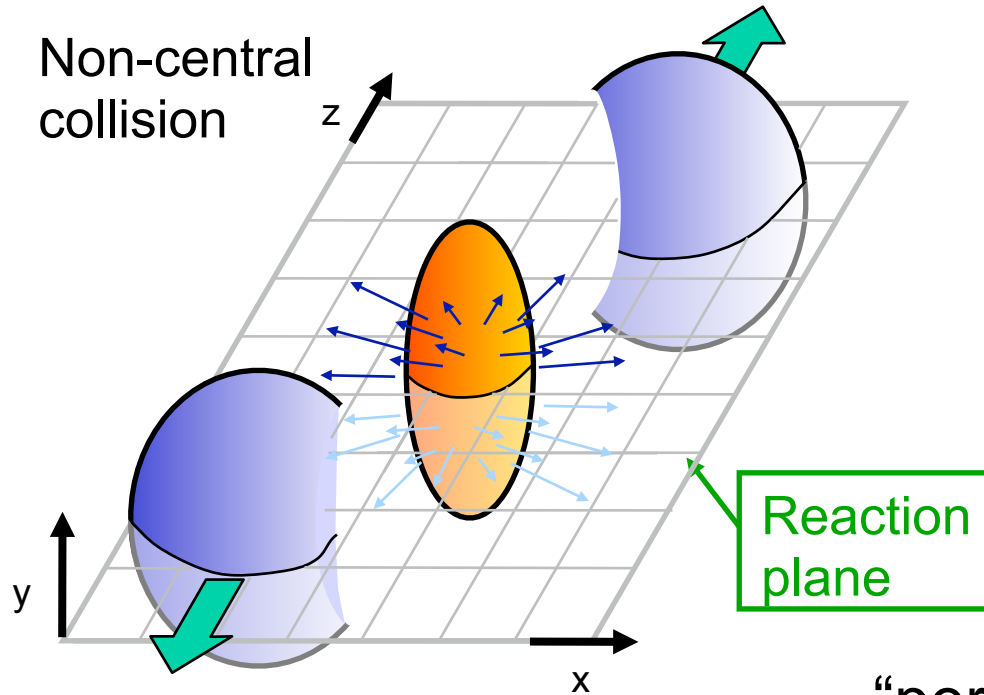


“peripheral” collision ($b \sim b_{\text{max}}$)

“central” collision ($b \sim 0$)

Geometry of a heavy-ion collision

Non-central collision



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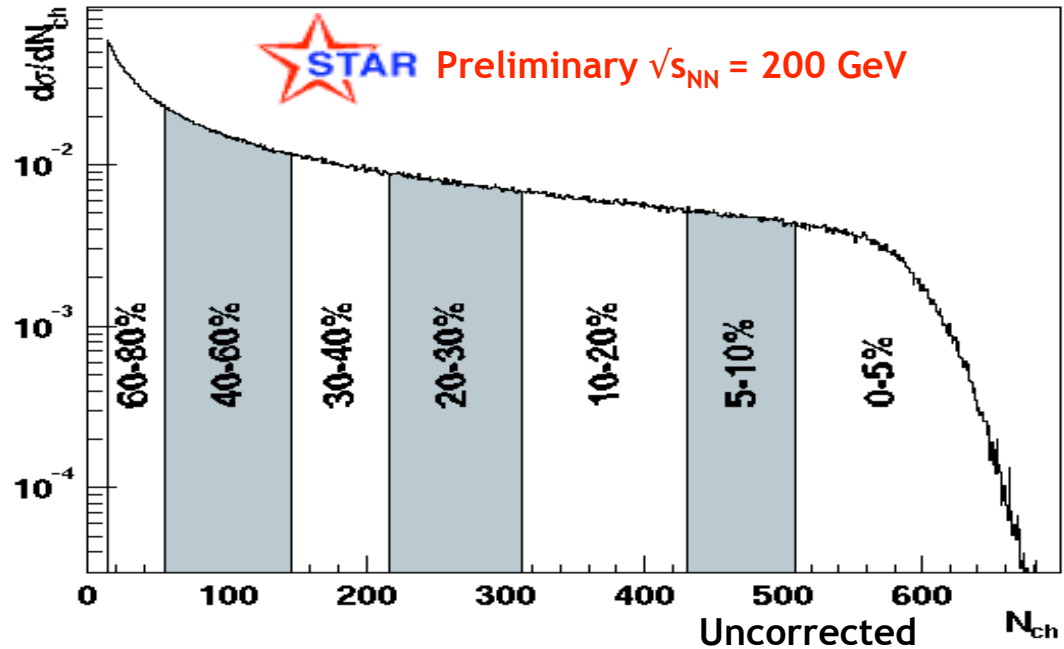
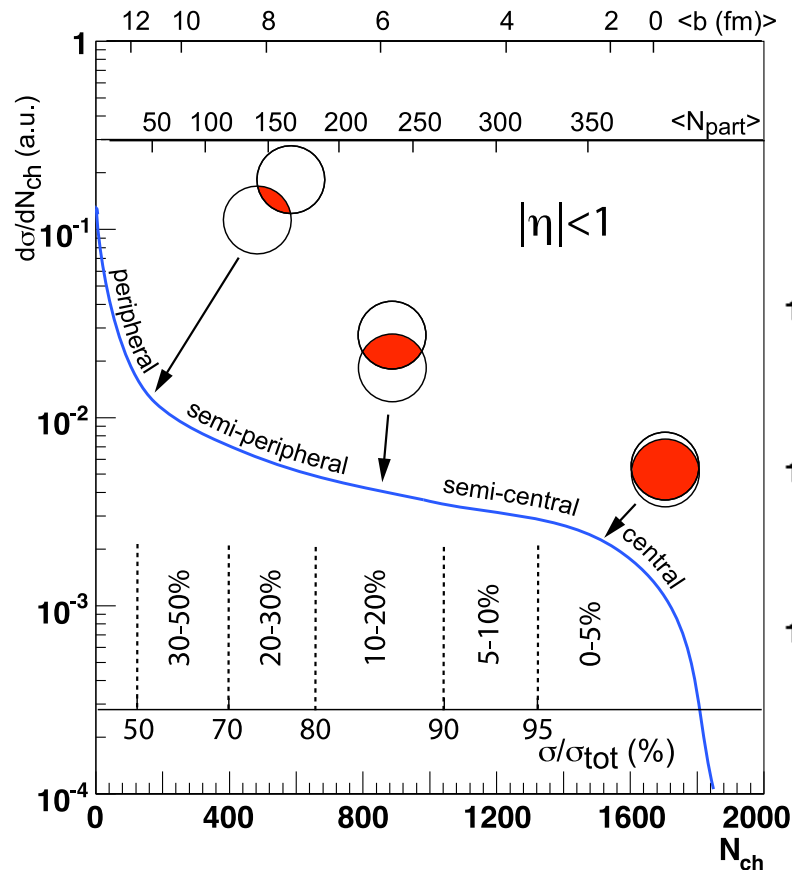
“central” collision ($b \sim 0$)

Number of participants (N_{part}): number of incoming nucleons (participants) in the overlap region

Number of binary collisions (N_{bin}): number of equivalent inelastic nucleon-nucleon collisions

$$N_{\text{bin}} \geq N_{\text{part}}/2$$

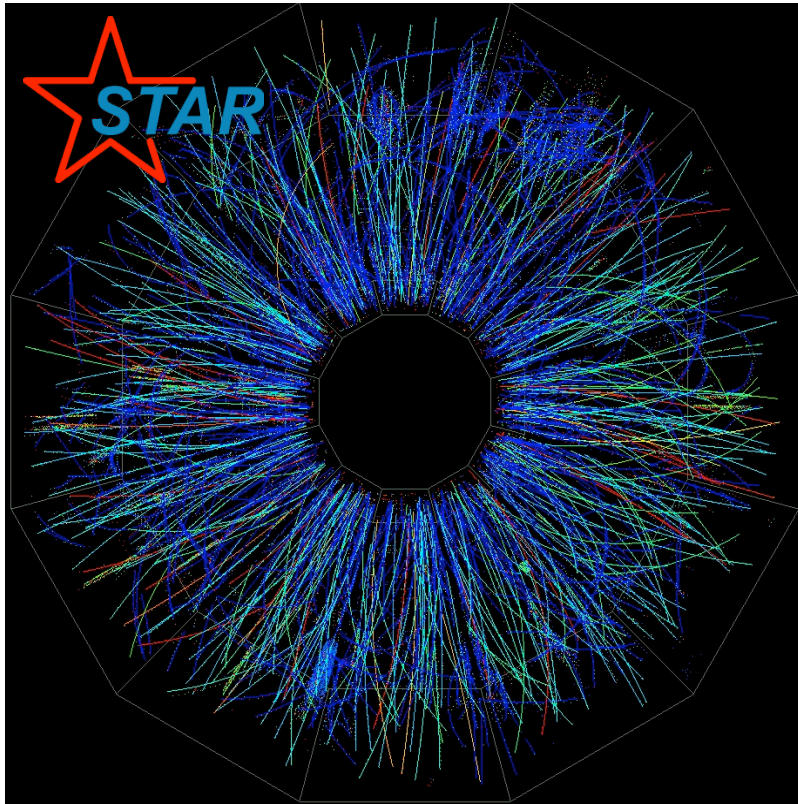
Comparing to data heavy-ion collision



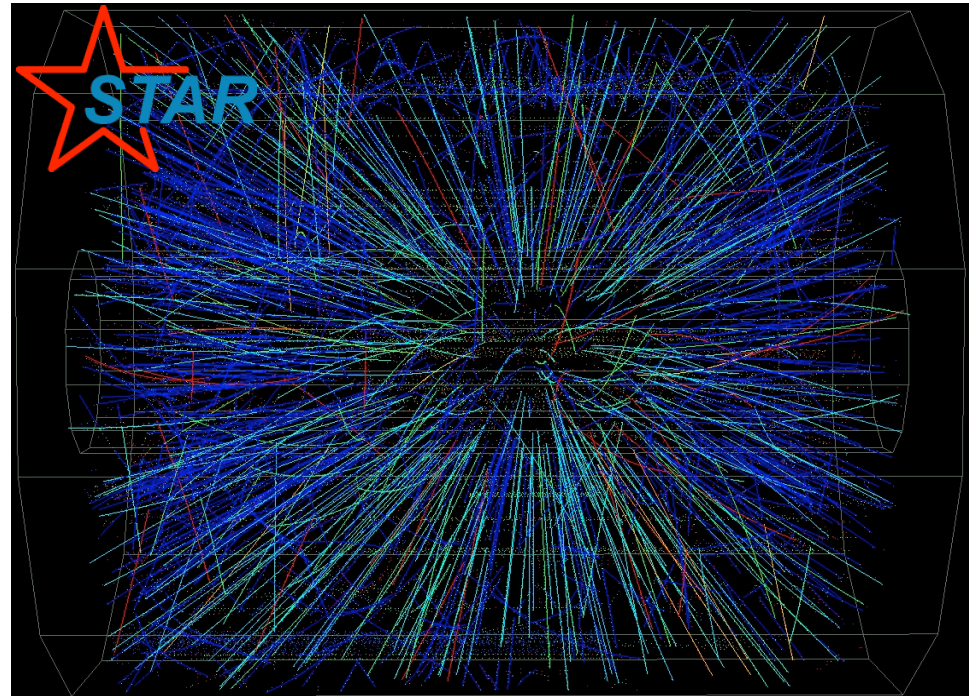
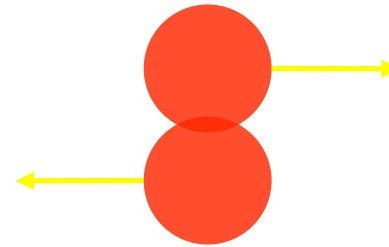
Good agreement between data and calculation

Measured mid-rapidity particle yield can be related to size of overlap region

A peripheral Au-Au collision

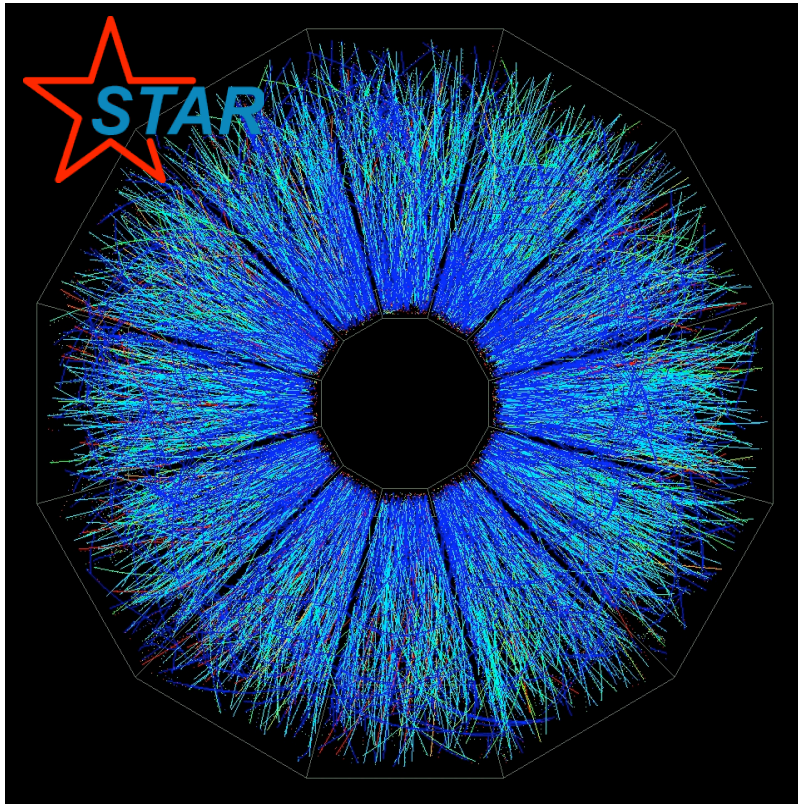


Peripheral Collision



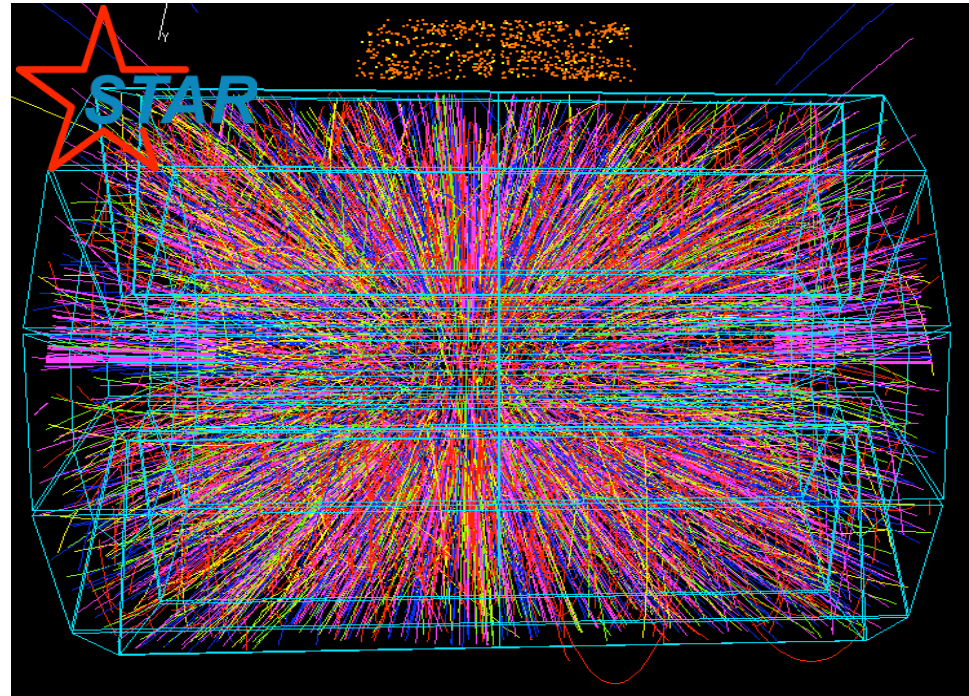
Color \Rightarrow Energy loss in TPC gas

39.4 TeV in central Au-Au collision

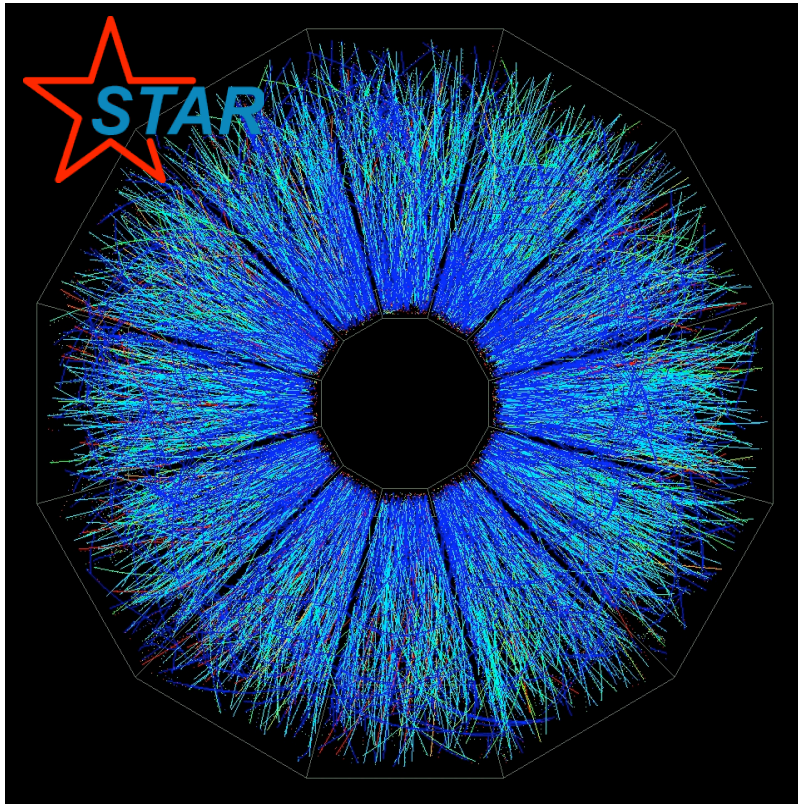


>5000 hadrons and leptons

- Only **charged** particles shown
- Neutrals don't ionise the TPC's gas so are not "seen" by this detector.



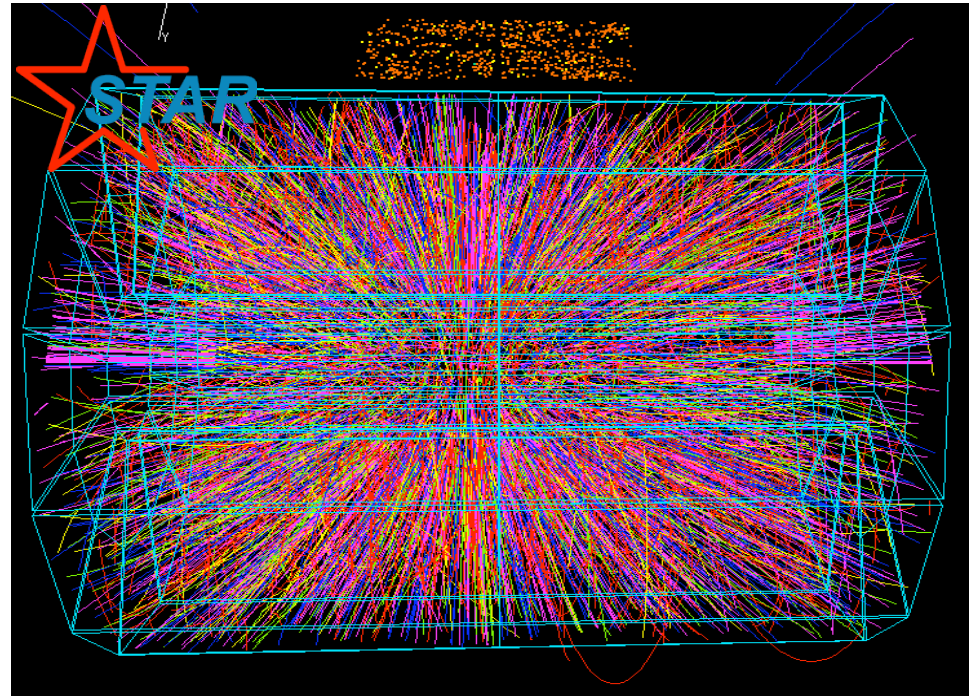
39.4 TeV in central Au-Au collision



>5000 hadrons and leptons

26 TeV is removed
from colliding beams.

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The energy is contained in one collision

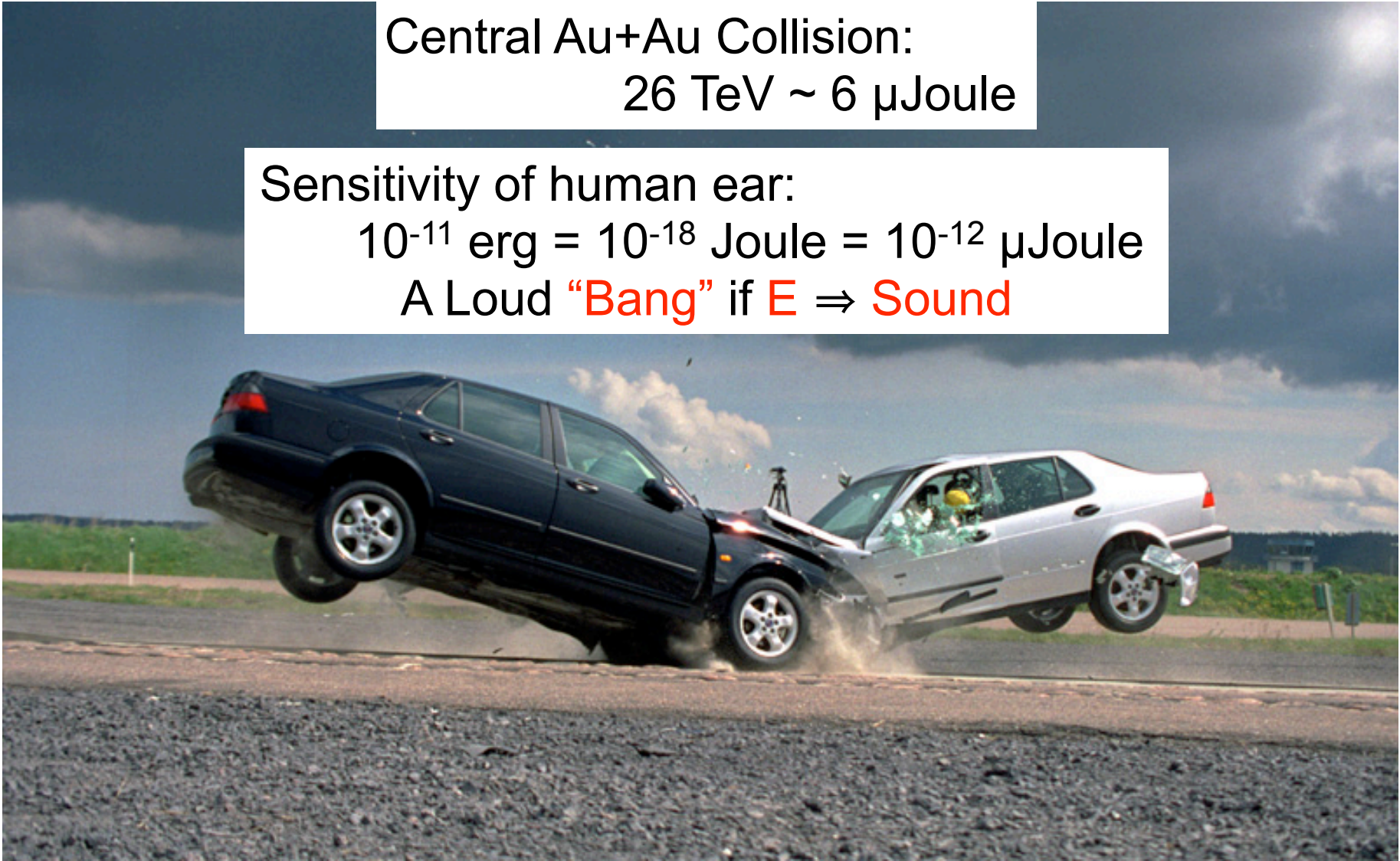
Central Au+Au Collision:
26 TeV \sim 6 μ Joule



The energy is contained in one collision

Central Au+Au Collision:
26 TeV \sim 6 μ Joule

Sensitivity of human ear:
 10^{-11} erg = 10^{-18} Joule = 10^{-12} μ Joule
A Loud “Bang” if $E \Rightarrow$ Sound



The energy is contained in one collision

Central Au+Au Collision:
26 TeV \sim 6 μ Joule

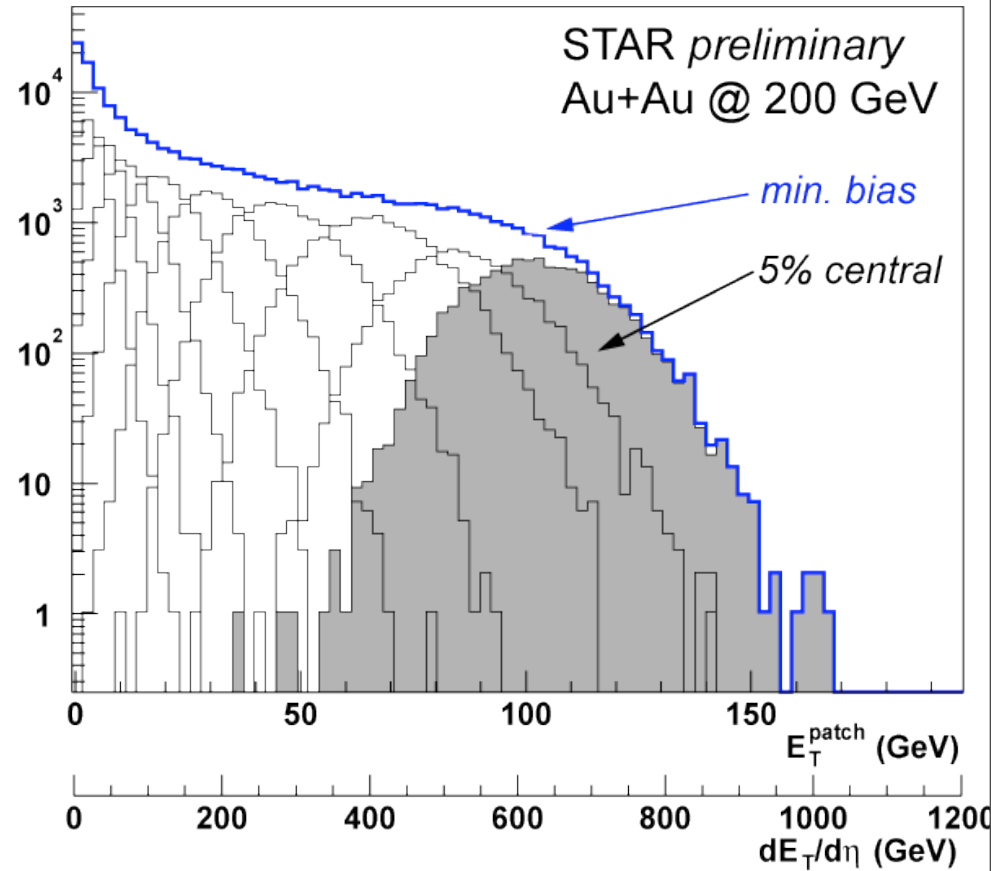
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Most goes into particle creation

Energy density in central Au-Au collisions

- use calorimeters to measure total energy



Energy density in central Au-Au collisions

- use calorimeters to measure total energy

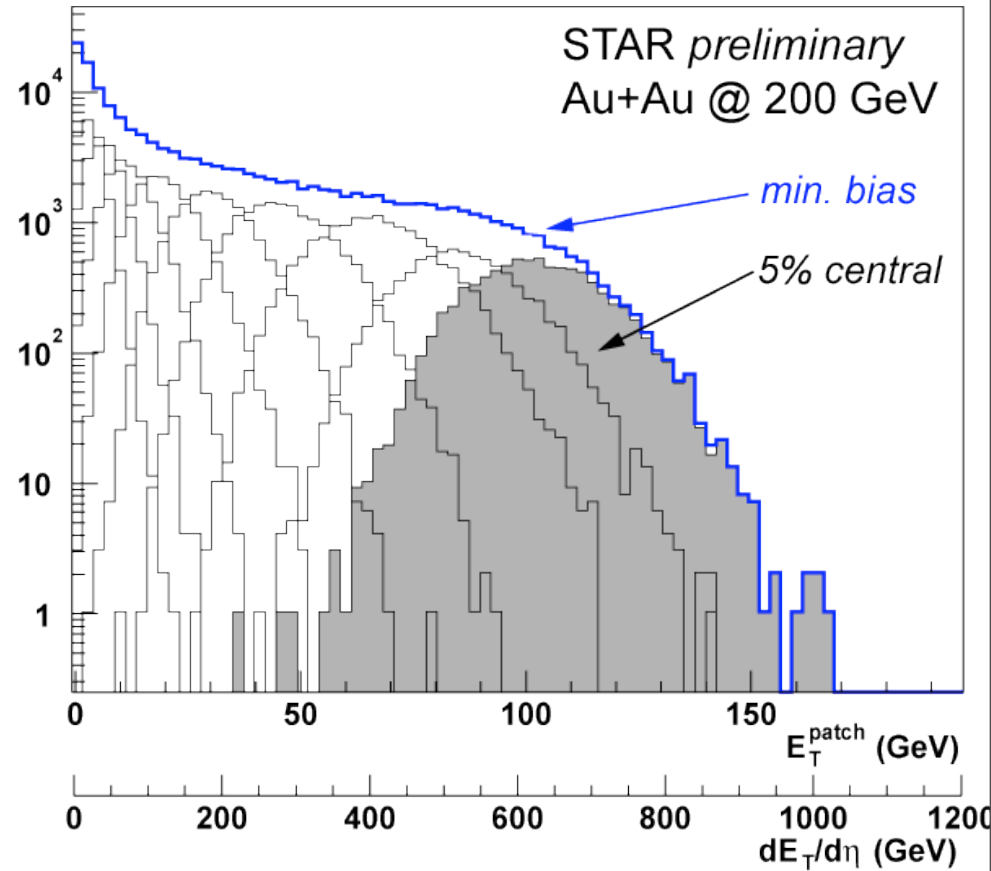
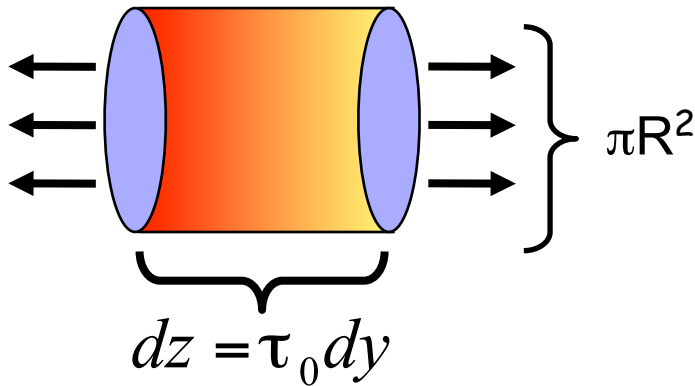
- estimate volume of collision

Bjorken-Formula for Energy Density:

$$\varepsilon_{Bj} = \frac{\Delta E_T}{\Delta V} = \frac{1}{\pi R^2} \frac{1}{\tau_0} \frac{dE_T}{dy}$$

$R \sim 6.5$ fm

Time it takes to thermalize system
($\tau_0 \sim 1$ fm/c)



Energy density in central Au-Au collisions

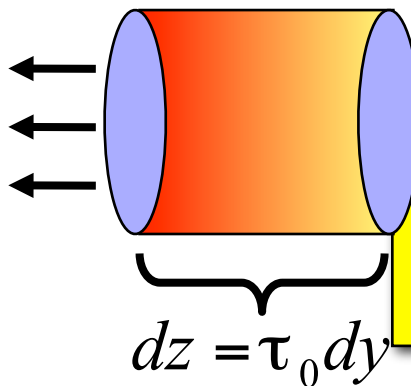
- use calorimeters to measure total energy
- estimate volume of collision

Bjorken-Formula for Energy Density:

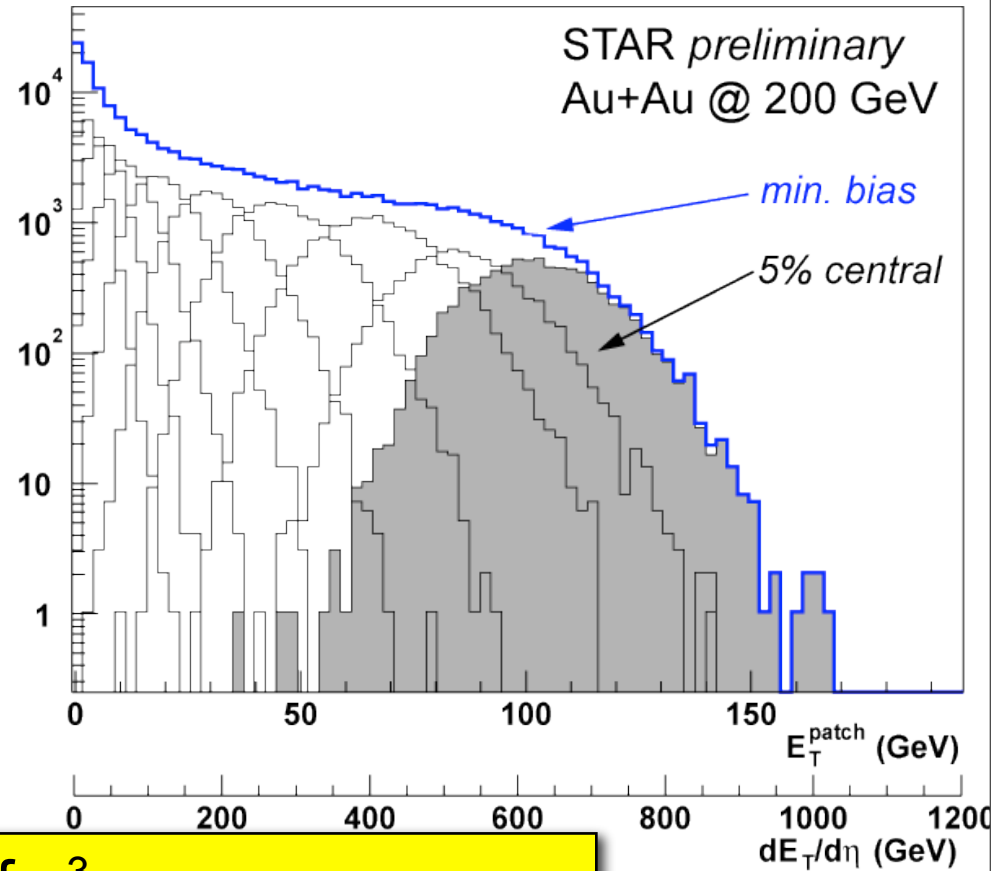
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$R \sim 6.5 \text{ fm}$

Time it takes to thermalize system
($\tau_0 \sim 1 \text{ fm}/c$)



$\epsilon_{BJ} \approx 5.0 \text{ GeV}/\text{fm}^3$
 ~ 30 times normal nuclear density
 ~ 5 times $> \epsilon_{\text{critical}}$ (lattice QCD)



5 GeV/fm³. Is that a lot?

In a year, the U.S. uses ~100 quadrillion BTUs of energy
(1 BTU = 1 burnt match):

$$100 \times 10^{15} BTU \times \frac{1060J}{BTU} \times \frac{1eV}{1.6 \times 10^{-19}J} = 6.6 \times 10^{38} eV$$

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At 5 GeV/fm³, this would fit in a volume of:

$$6.6 \times 10^{38} eV \div \frac{5 \times 10^9 eV}{fm^3} = 1.3 \times 10^{29} fm^3$$

5 GeV/fm³. Is that a lot?

In a year, the U.S. uses ~100 quadrillion BTUs of energy
(1 BTU = 1 burnt match):

$$100 \times 10^{15} BTU \times \frac{1060J}{BTU} \times \frac{1eV}{1.6 \times 10^{-19}J} = 6.6 \times 10^{38} eV$$

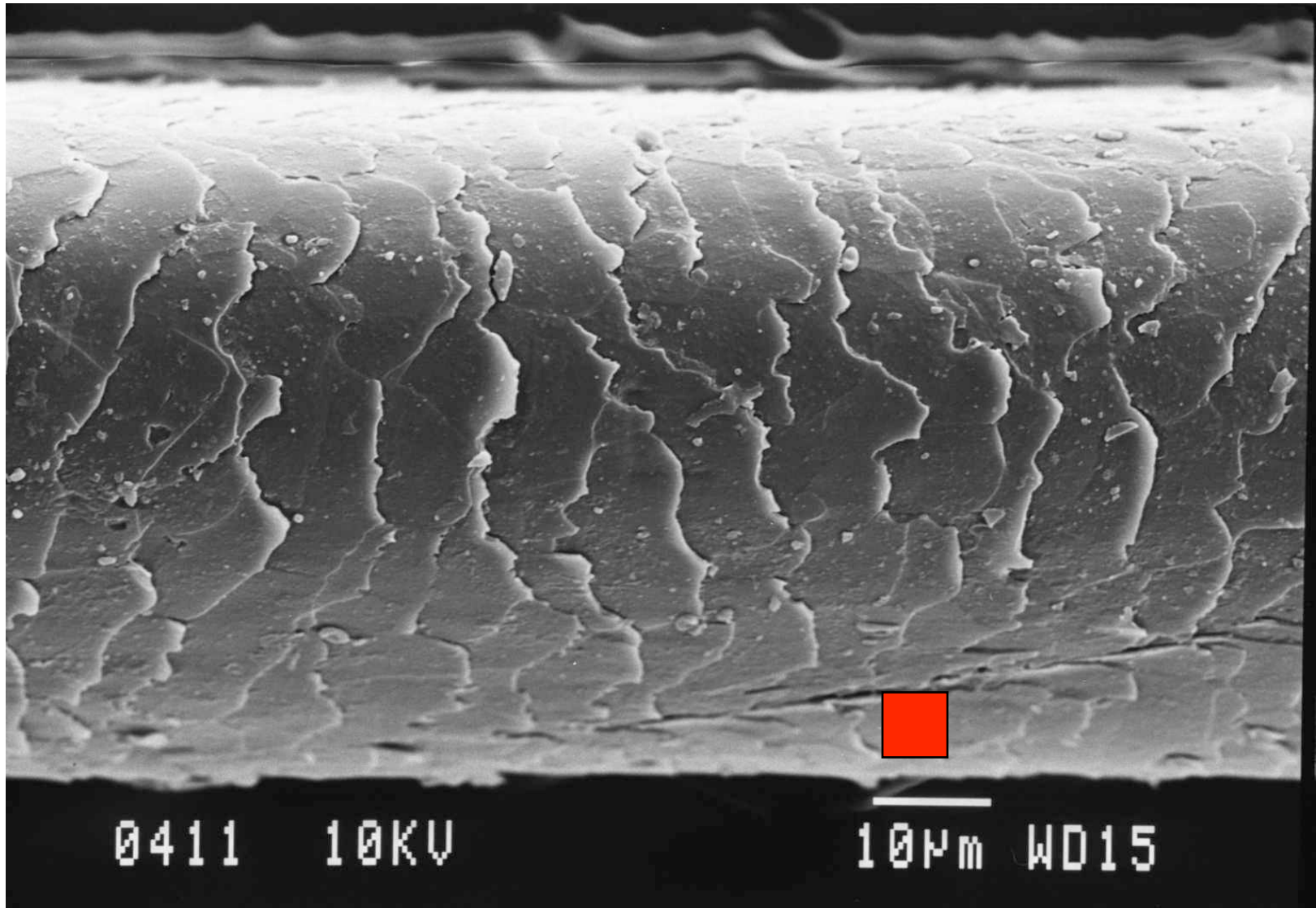
At 5 GeV/fm³, this would fit in a volume of:

$$6.6 \times 10^{38} eV \div \frac{5 \times 10^9 eV}{fm^3} = 1.3 \times 10^{29} fm^3$$

Or, in other words, in a box of the following dimensions:

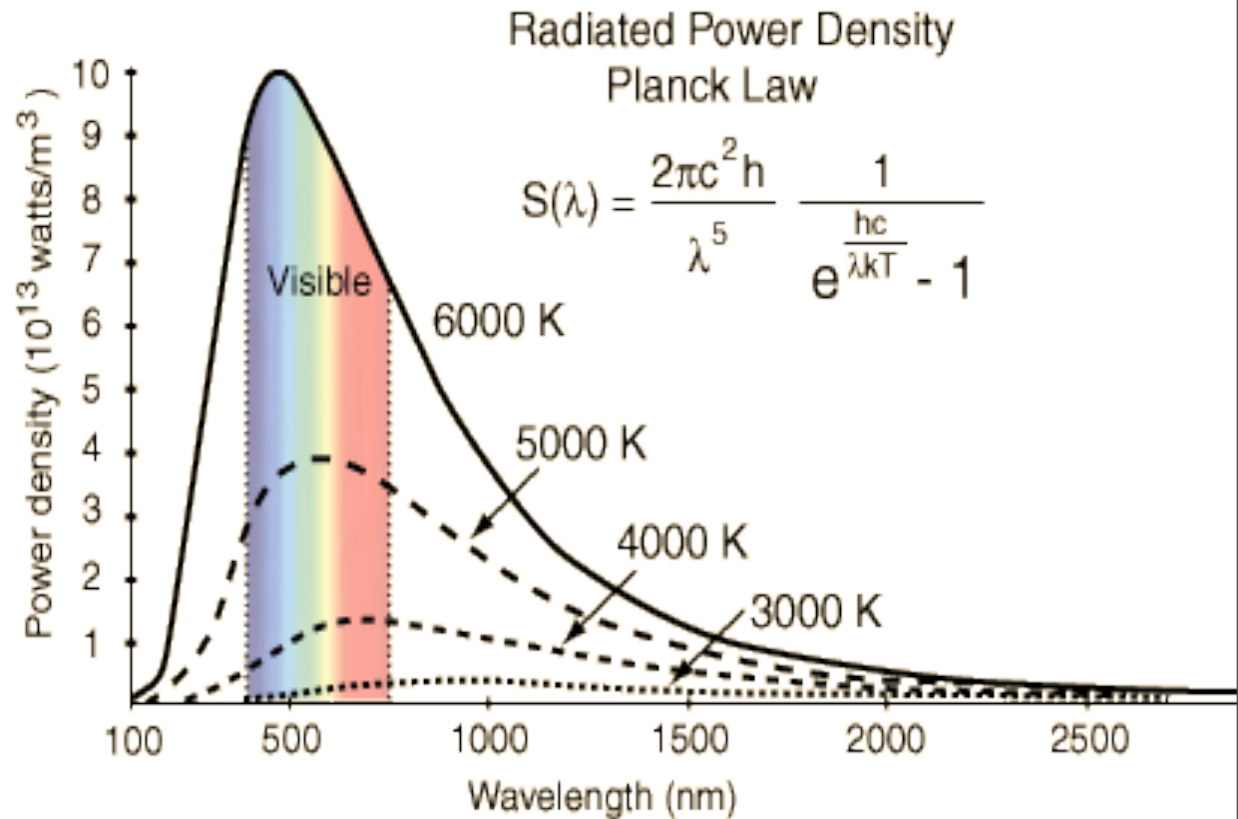
$$\sqrt[3]{1.3 \times 10^{29} fm^3} = 5 \times 10^9 fm = 5 \mu m$$

A human hair



Measuring the initial temperature

Planck distribution describes **intensity** as a **function of the wavelength** of the emitted radiation



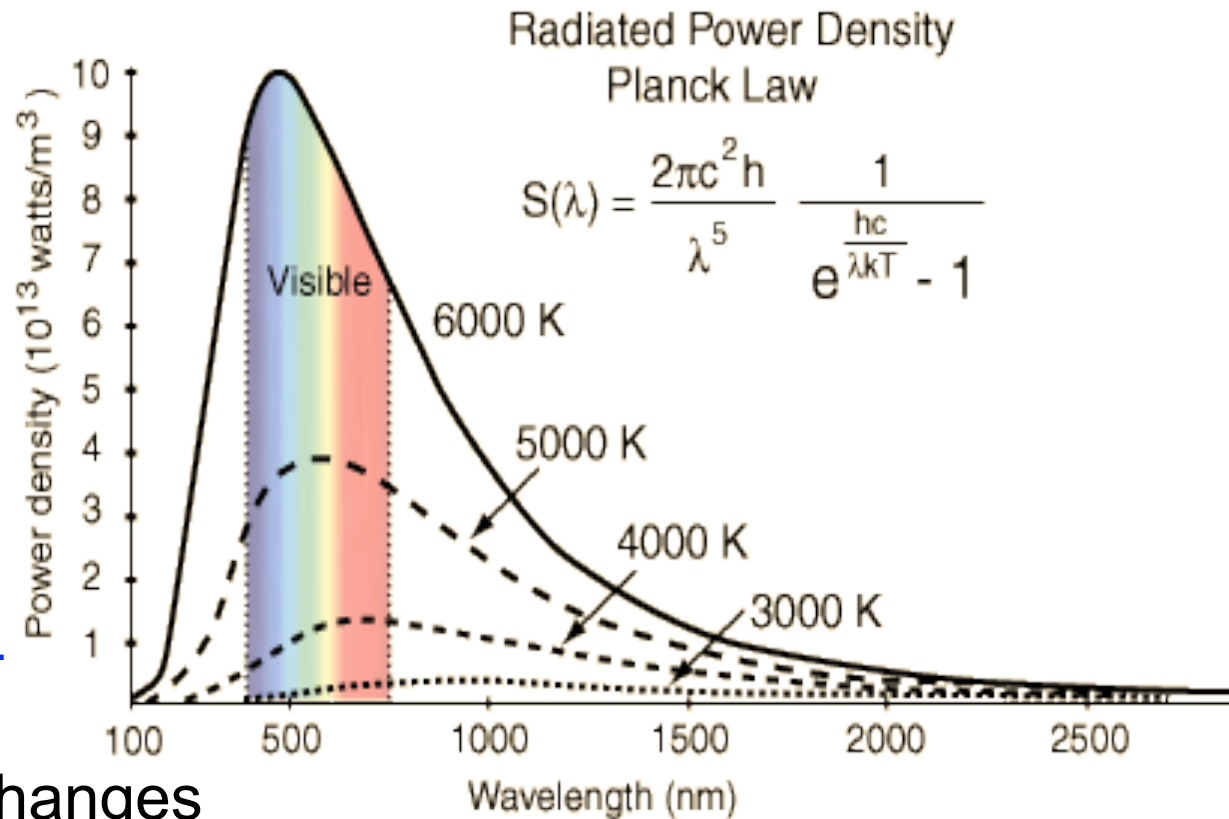
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“Blackbody” radiation is the spectrum of radiation emitted by an object at temperature T

As T increases curve changes

Photons have no charge or color → don't interact with medium

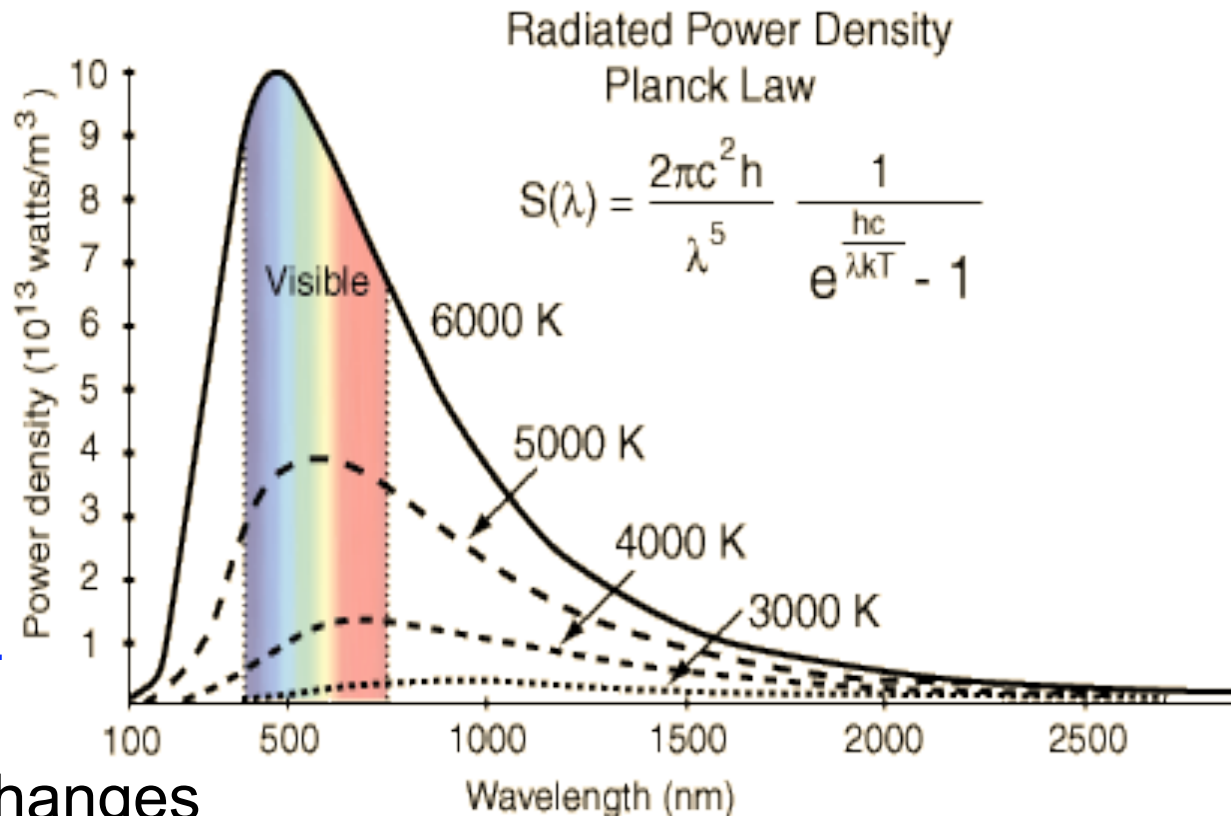


Measuring the initial temperature

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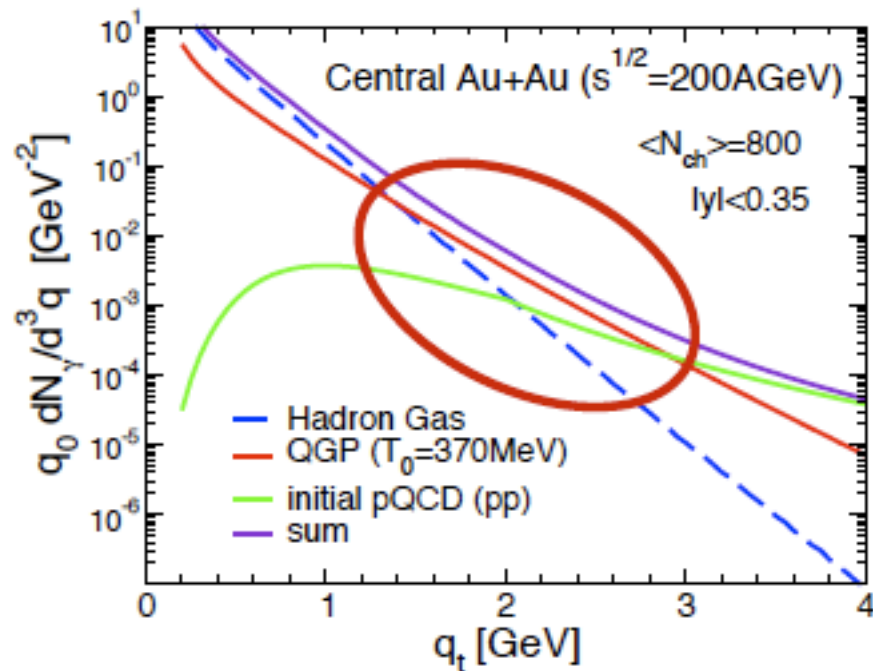
γ reveals early temperature of QGP

Measuring the initial temperature

γ emitted over all lifetime \rightarrow convolution of all temperatures

Theory now well developed

Corrections under control



QGP dominates
 $1 < p_T < 3 \text{ GeV}/c$

Experimentally photons a
complete pain!

Calorimetry

granularity and resolution

Large background

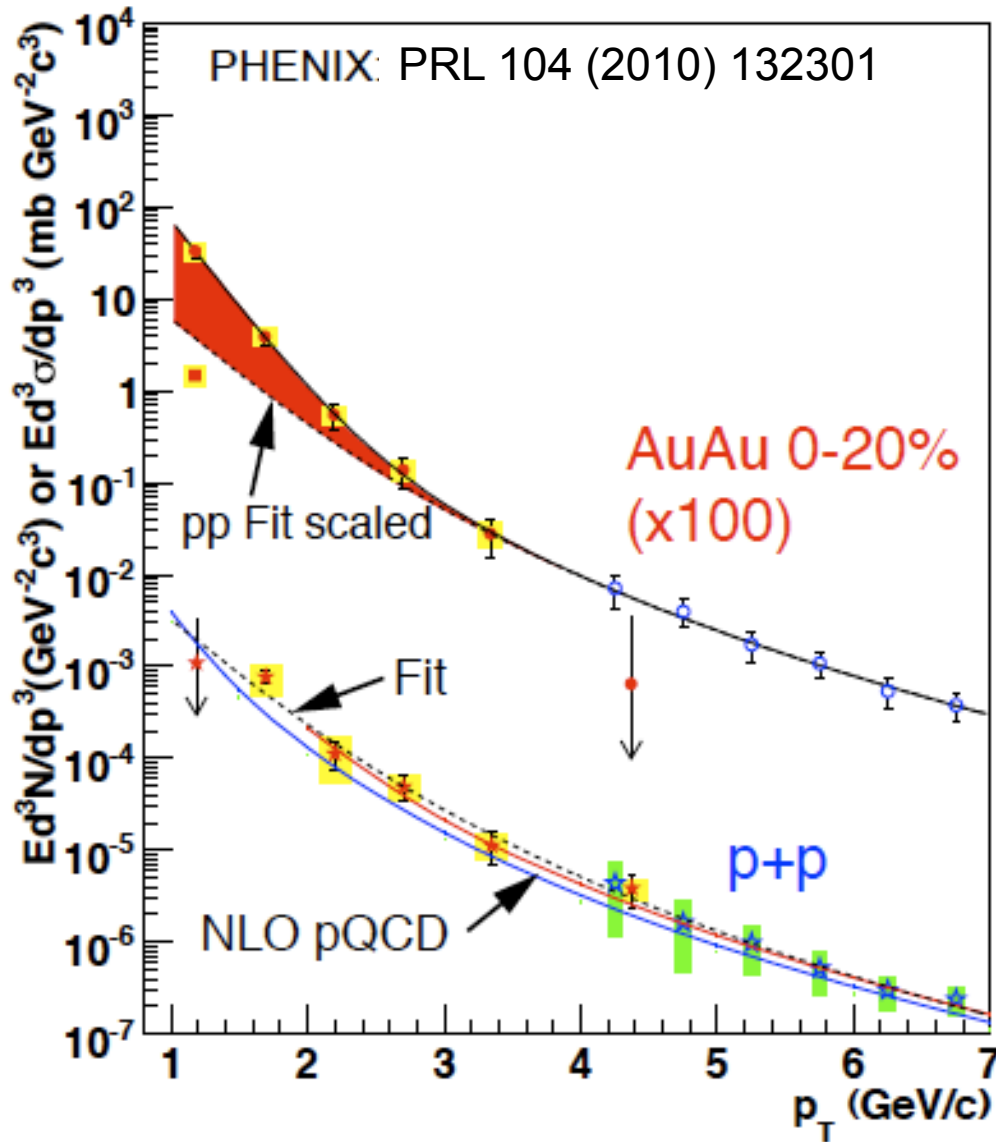
$\pi^0/\eta \rightarrow \gamma\gamma$

Losses from conversions

Hadronic contamination

Signal:Background $\sim 1:10$

Measuring the initial temperature



After background subtraction:

Emission rate and shape
consistent with that from a
hot thermally equilibrated
medium

$T = 300 - 600 \text{ MeV}$
 $\tau = 0.15 - 0.6 \text{ fm/c}$

Large uncertainty due to correlated pair background
i.e. jets

What is the temperature of the medium?

- Statistical Thermal Models:
 - Assume a system that is **thermally** (constant T_{ch}) and **chemically** (constant n_i) **equilibrated**
 - System composed of non-interacting hadrons and resonances
 - Obey conservation laws: Baryon Number, Strangeness, Isospin
- Given T_{ch} and μ 's (+ system size), n_i 's can be calculated in a grand canonical ensemble

$$n_i = \frac{g}{2\pi^2} \int_0^\infty \frac{p^2 dp}{e^{(E_i(p) - \mu_i)/T} \pm 1}, \quad E_i = \sqrt{p^2 + m_i^2}$$

Fitting the particle ratios

Number of particles of a given species related to temperature

$$dn_i \sim e^{-(E-\mu_B)/T} d^3p$$

- Assume all particles described by same temperature T and μ_B
- one ratio (e.g., \bar{p} / p) determines μ / T :

$$\frac{\bar{p}}{p} = \frac{e^{-(E+\mu_B)/T}}{e^{-(E-\mu_B)/T}} = e^{-2\mu_B/T}$$

- A second ratio (e.g., K / π) provides $T \rightarrow \mu$

$$\frac{K}{\pi} = \frac{e^{-E_K/T}}{e^{-E_\pi/T}} = e^{-(E_K-E_\pi)/T}$$

- Then all other hadronic ratios (and yields) defined

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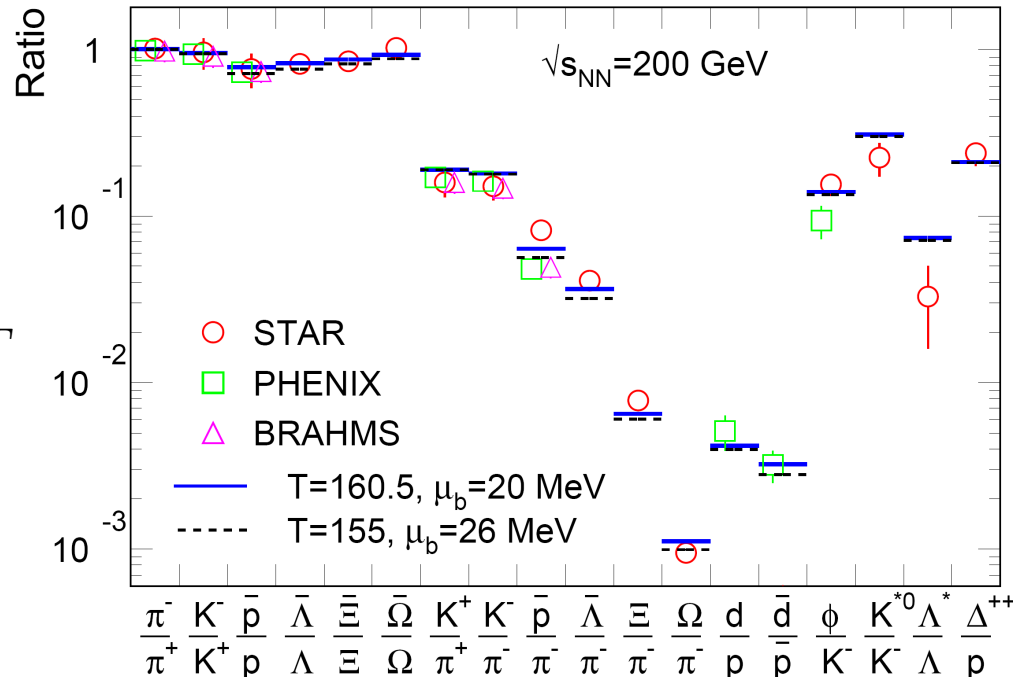
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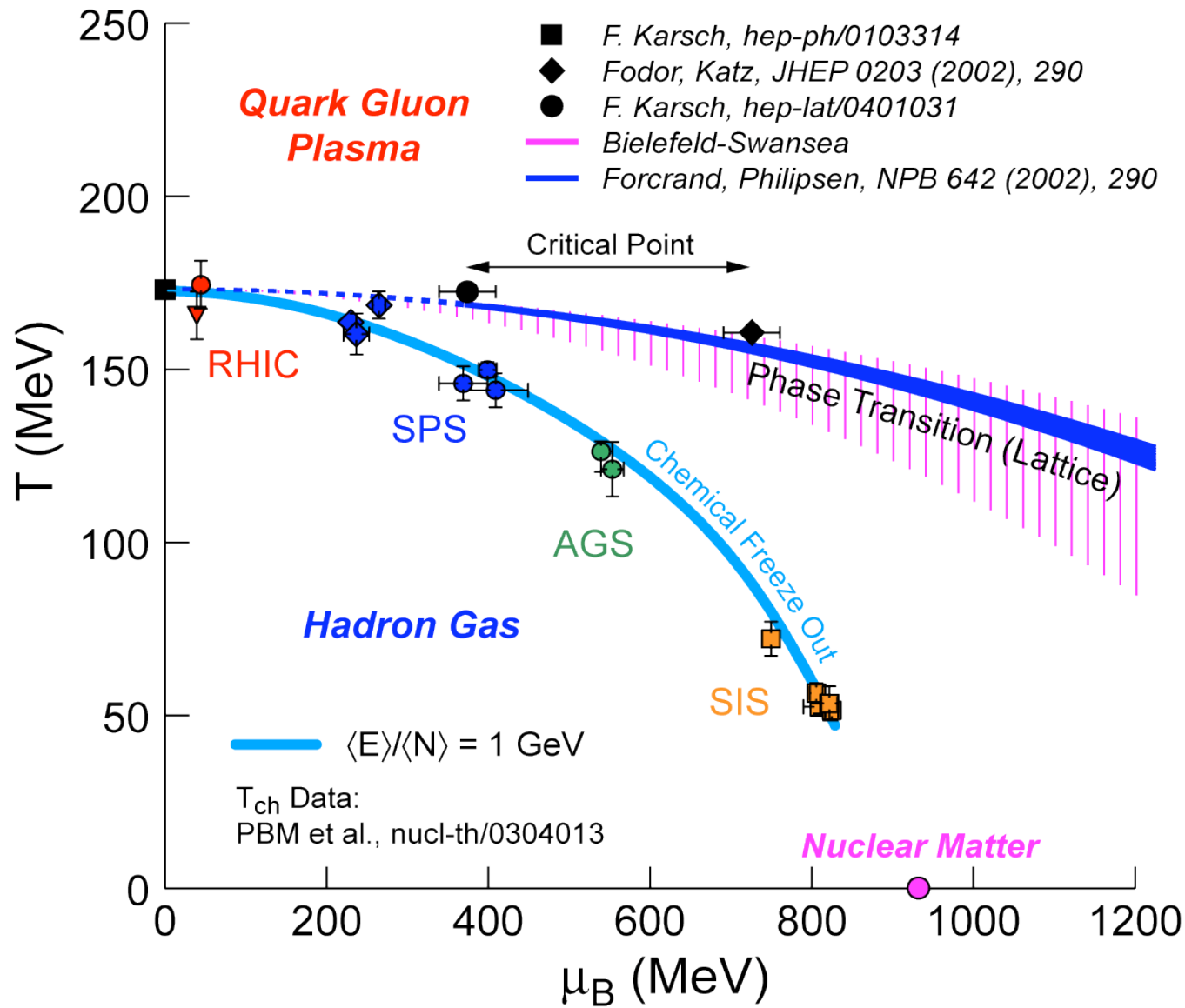
A. Adronic *et al.*, NPA772:167



$T \sim 160 \text{ MeV}, \mu_b \sim 20 \text{ MeV}$

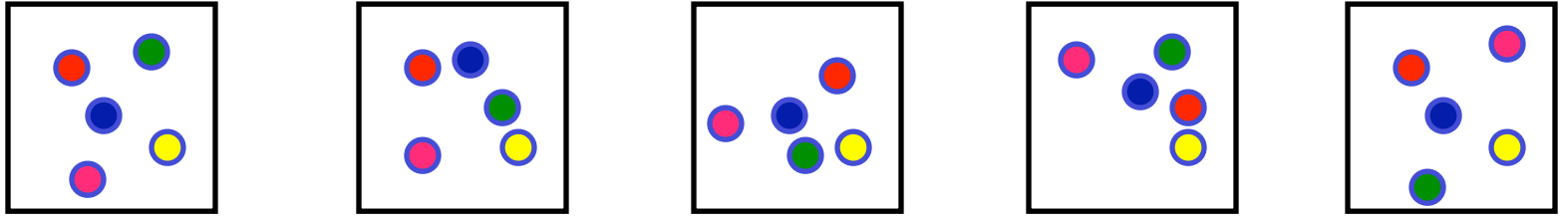
Initial Temperature
probably much higher

Where RHIC sits on the phase diagram



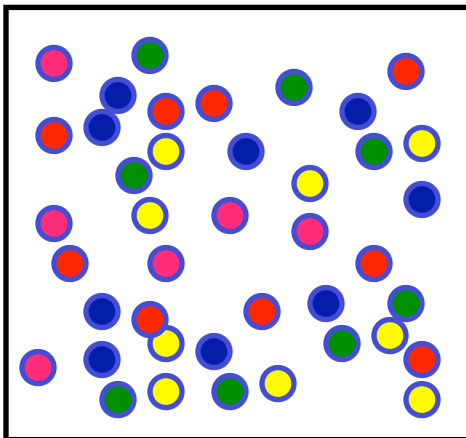
Statistics \neq thermodynamics

$p+p$



Ensemble of events constitutes a statistical ensemble
 T and μ are simply Lagrange multipliers
“Phase Space Dominance”

$A+A$



One (1) system is already statistical !

- We can talk about pressure
- T and μ are more than Lagrange multipliers

By the way!

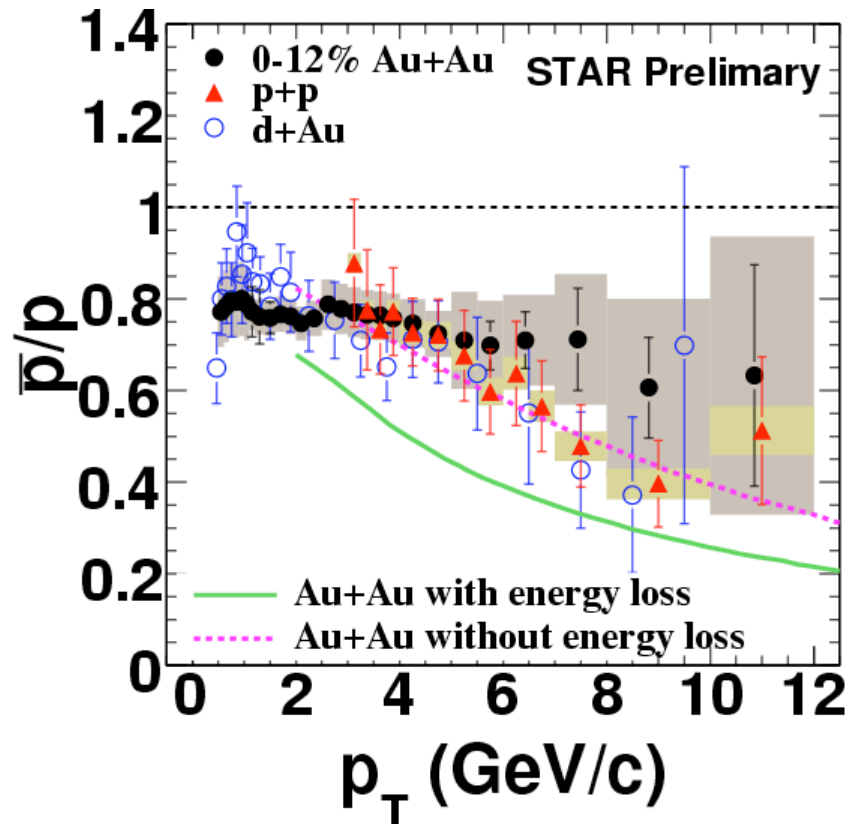
Take a second look at the anti-proton/proton ratio

$$\bar{p}/p \sim 0.8$$

There is a net baryon number at mid-rapidity!!

Baryons number is being transported over 6 units of rapidity from the incoming beams to the collision zone!

Baryon number not carried by quarks



- baryon junctions postulated

Evidence for thermalization?

- Not all processes which lead to multi-particle production are thermal - elementary collisions
- *Any mechanism for producing hadrons which evenly populates the free particle phase space will mimic a microcanonical ensemble.*
- Relative probability to find n particles is the ratio of the phase-space volumes $P_n/P_{n'} = \varphi_n(E)/\varphi_{n'}(E) \Rightarrow$ given by statistics only.
- *Difference between MCE and CE vanishes as the size of the system N increases.*
- Such a system is NOT in thermal equilibrium - to thermalize need interactions/re-scattering

Need to look for other evidence of collective motion

Determining the temperature

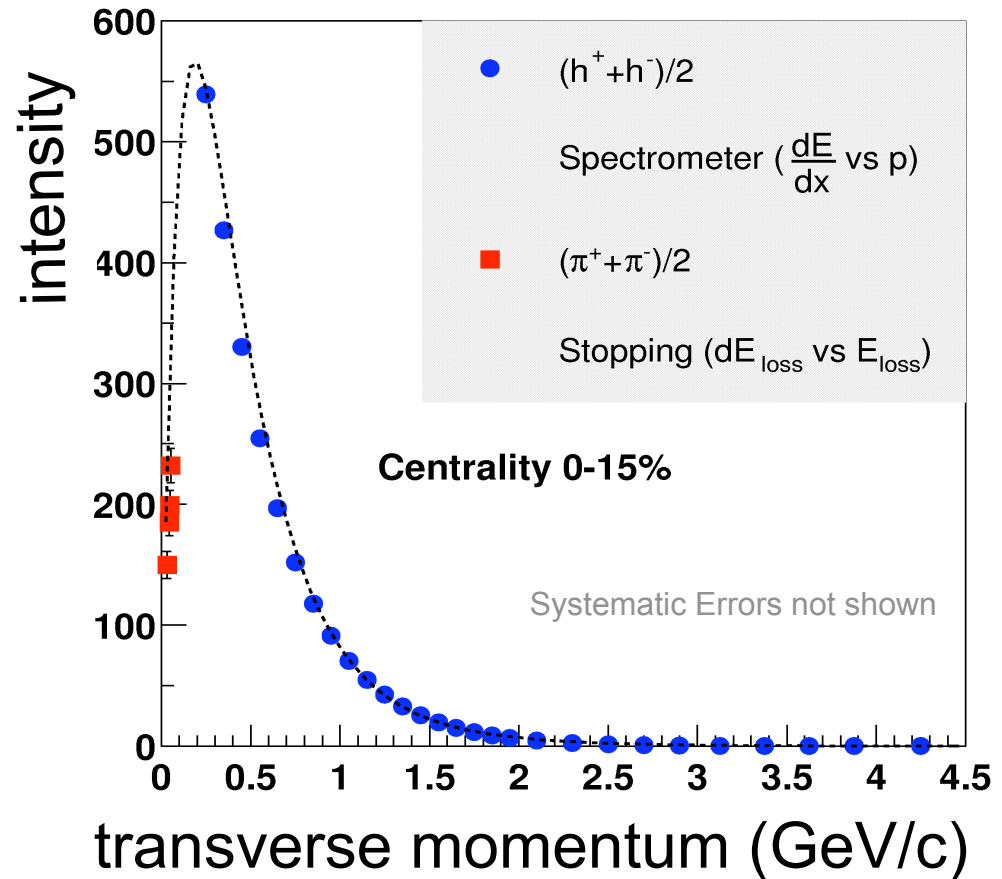
$$1/\text{Wavelength} \propto \text{Freq} \propto E \propto p$$

From transverse
momentum distribution of
pions deduce
temperature $\sim 120 \text{ MeV}$

$$E = \frac{3}{2}kT$$

$$T = \frac{2E}{3k}$$

$$= \frac{2 \times 120 \times 10^6}{3 \times 1.4 \times 10^{-23}} \times 1.6 \times 10^{-19}$$
$$\sim 9 \times 10^{11} K$$



Determining the temperature

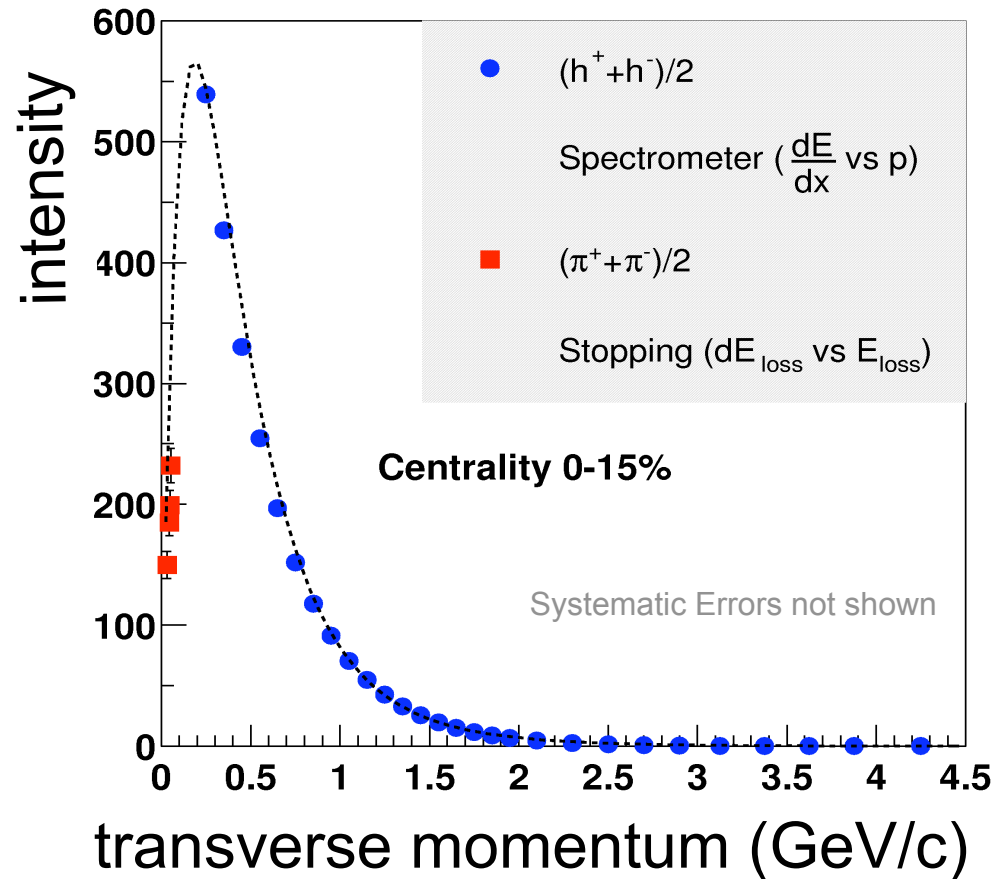
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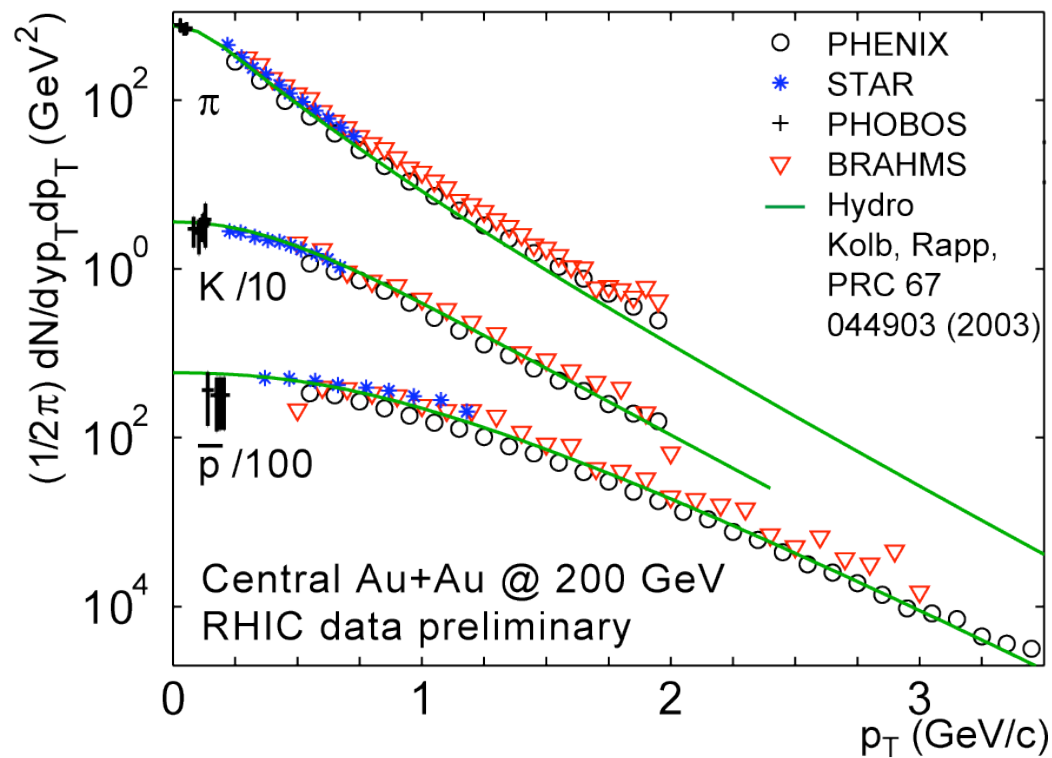
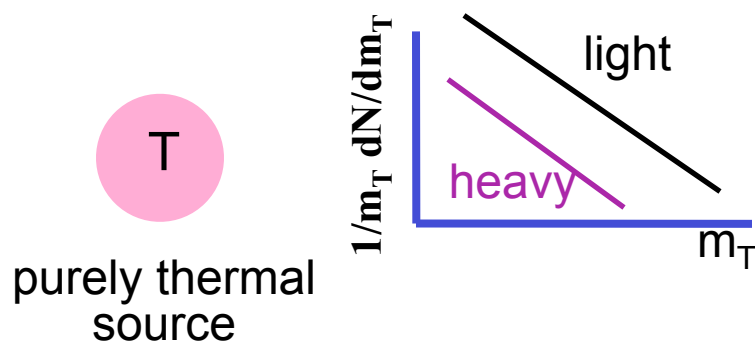
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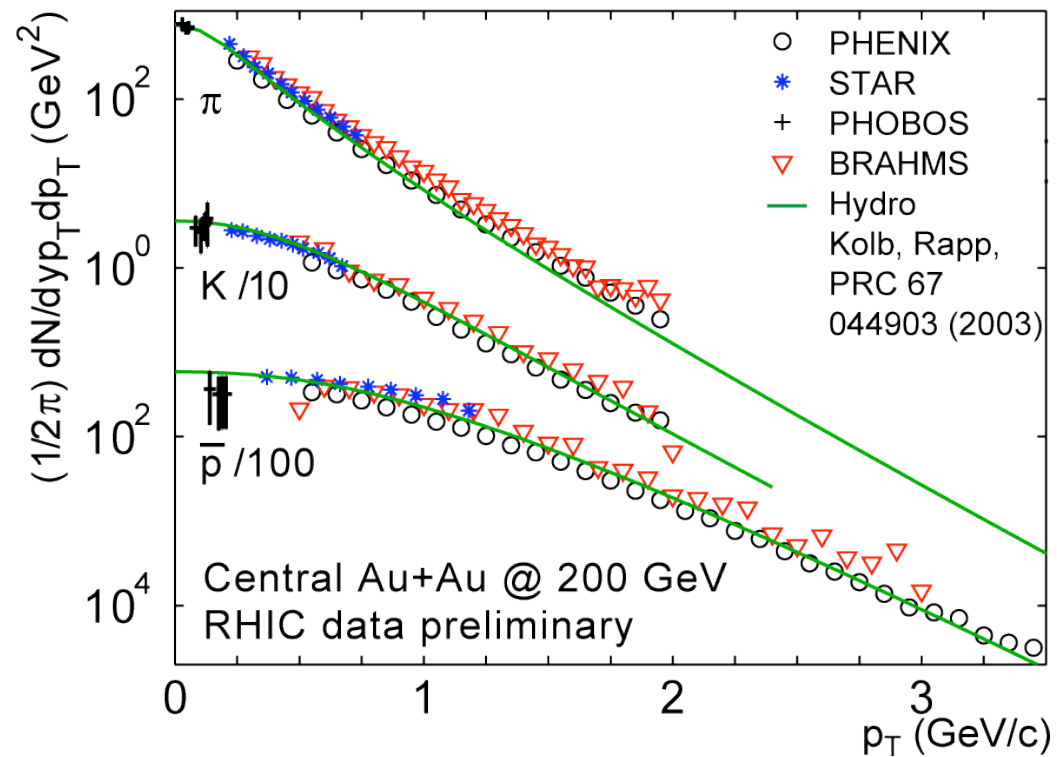
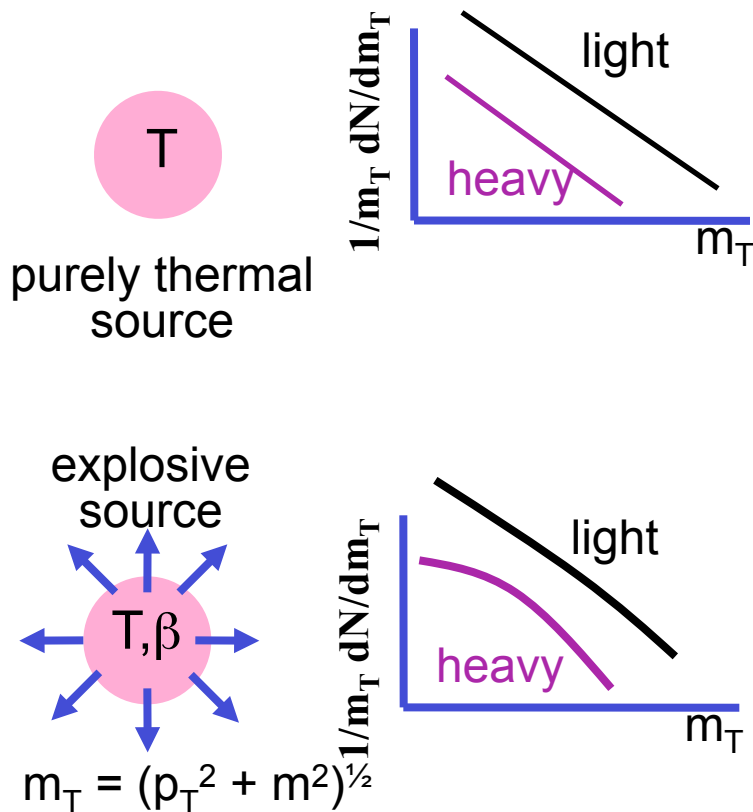
$$T_{\text{ch}} > T_{\text{fo}}$$

System exist for time in hadronic phase

Strong collective radial expansion

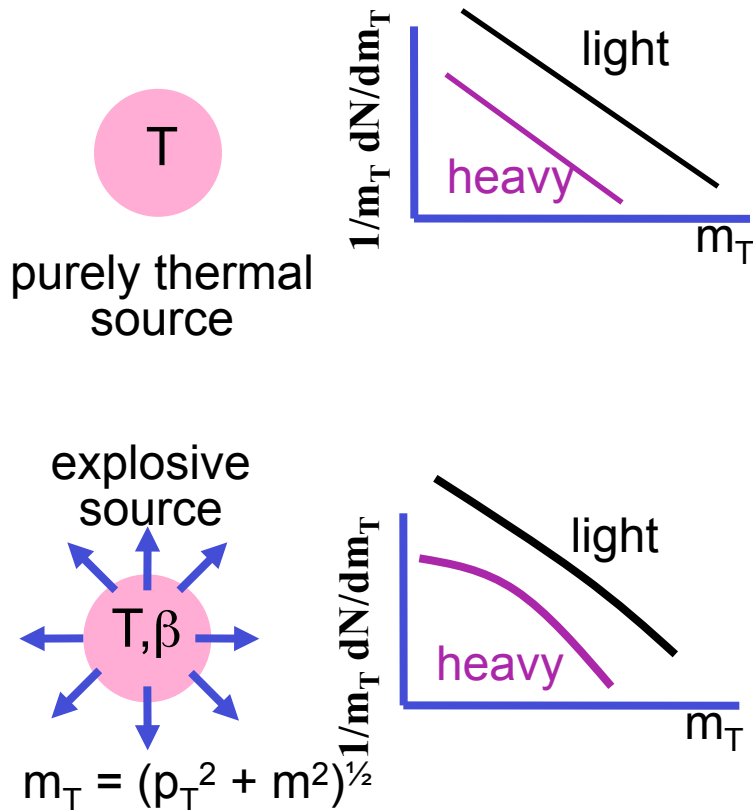


Strong collective radial expansion

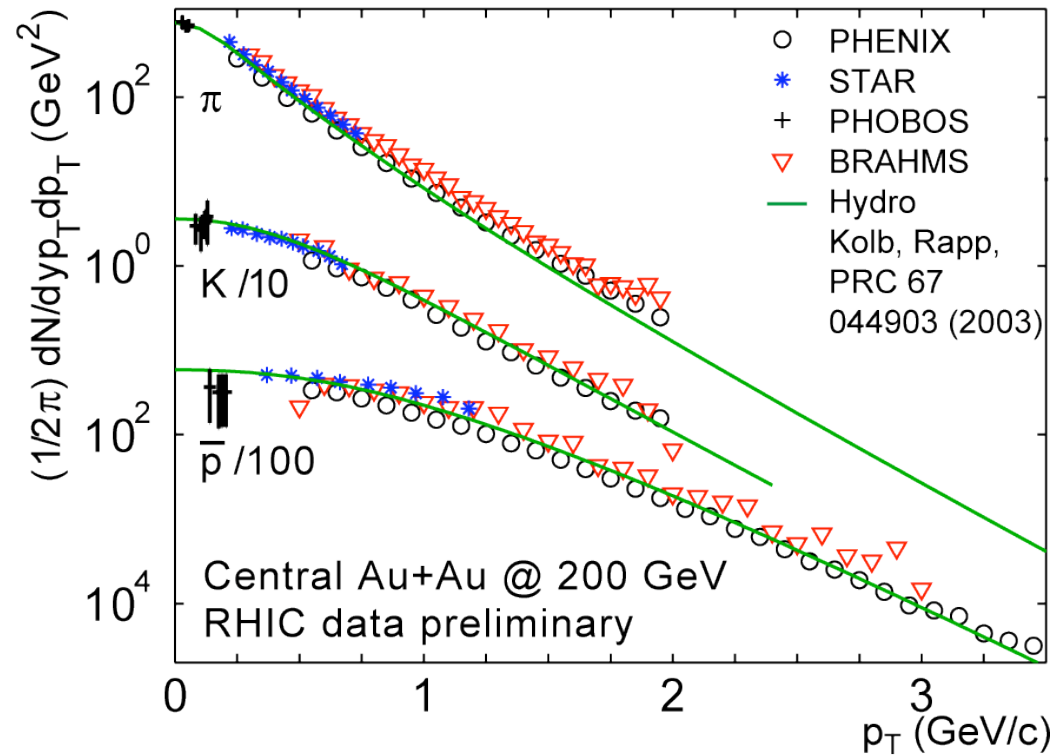


- Different spectral shapes for particles of differing mass
→ strong collective radial flow

Strong collective radial expansion



- Different spectral shapes for particles of differing mass
→ strong collective radial flow

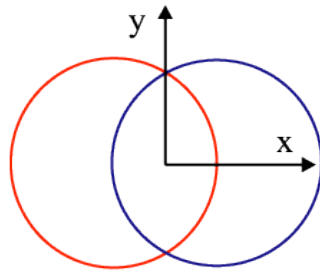


$$T_{fo} \sim 100 \text{ MeV}$$

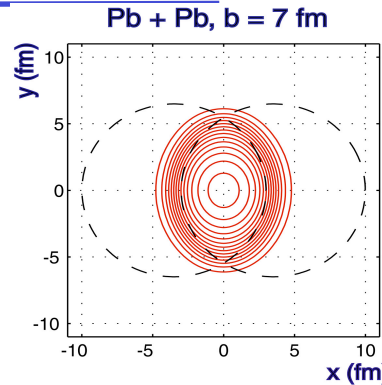
$$\langle \beta_T \rangle \sim 0.55 c$$

Good agreement with hydrodynamic prediction for soft EOS (QGP+HG)

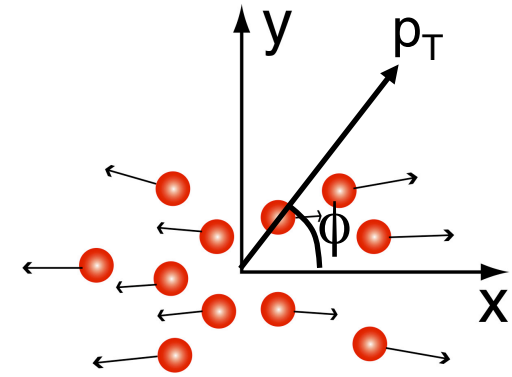
Anisotropic/Elliptic flow



Almond shape overlap
region in **coordinate space**



Interactions/
Rescattering



Anisotropy in
momentum space

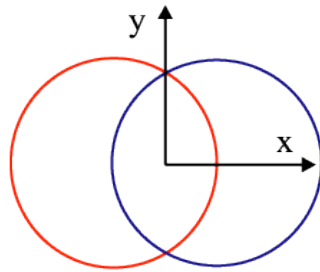
$$dN/d\phi \sim 1 + 2 v_2(p_T) \cos(2\phi) + \dots$$

$$\phi = \text{atan}(p_y/p_x)$$

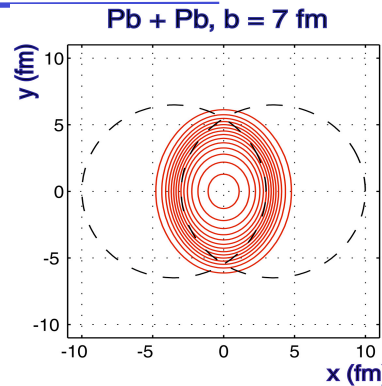
$$v_2 = \langle \cos 2\phi \rangle$$

v_2 : 2nd harmonic Fourier coefficient in $dN/d\phi$ with respect to the reaction plane

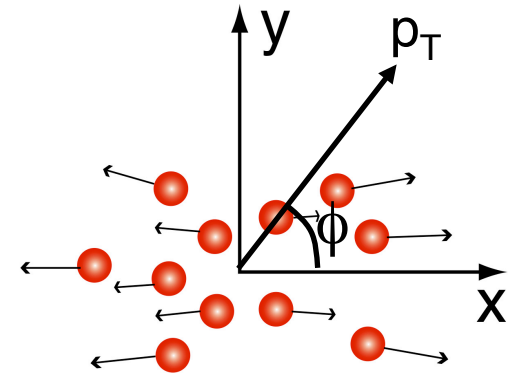
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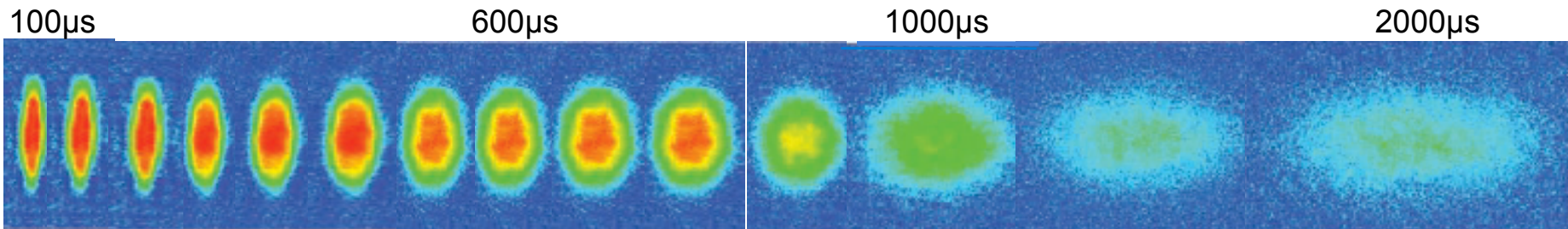
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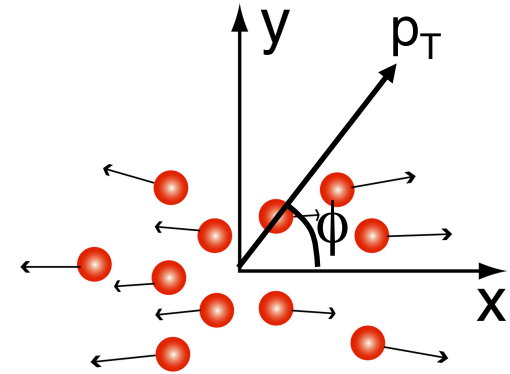
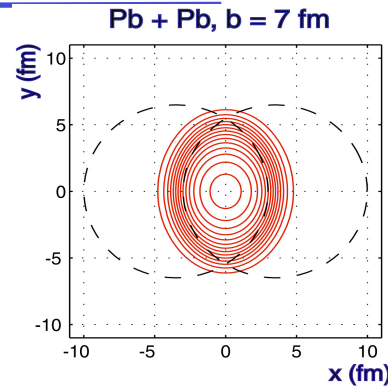
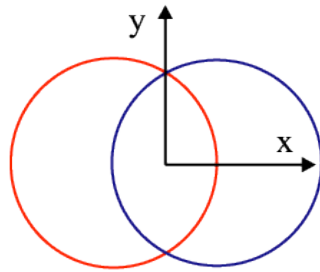
Time

—M. Gehm, S. Granade, S. Hemmer, K. O'Hara, J. Thomas - **Science** 298 **2179** (2002)

Helen Caines - HPCSS - August 2010

33

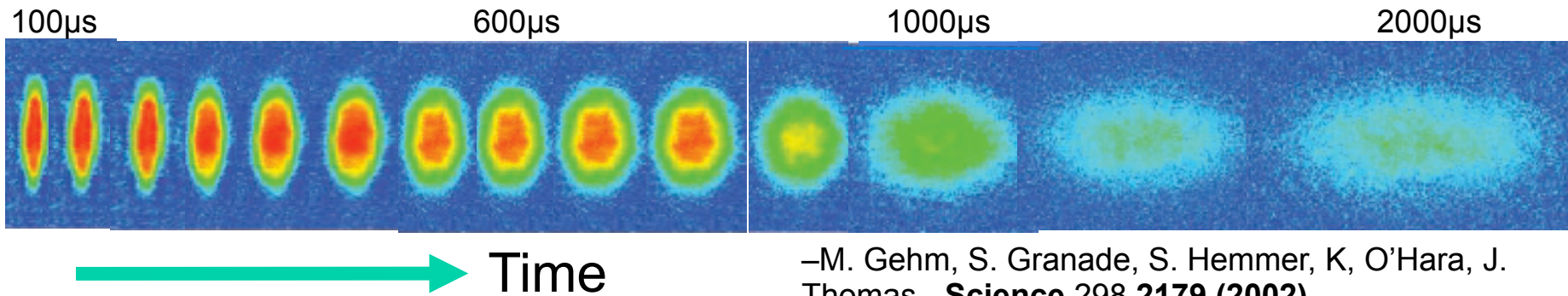
Anisotropic/Elliptic flow



Elliptic flow observable sensitive to early evolution of system

Mechanism is self-quenching

Large v_2 is an indication of **early** thermalization



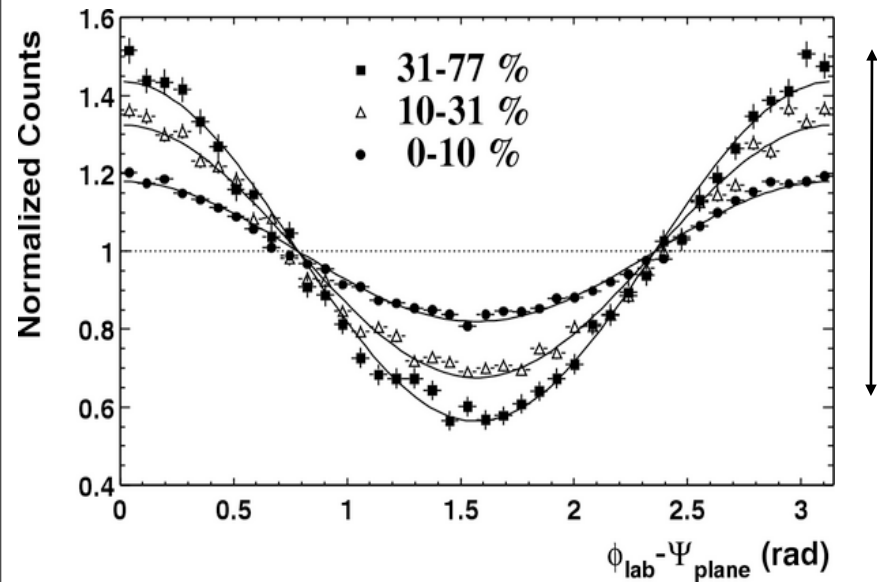
—M. Gehm, S. Granade, S. Hemmer, K. O'Hara, J. Thomas - **Science** 298 **2179** (2002)

Helen Caines - HPCSS - August 2010

33

Elliptic flow

Distribution of particles with respect to event plane, $\phi-\psi$, $p_t > 2$ GeV; STAR PRL 90 (2003) 032301

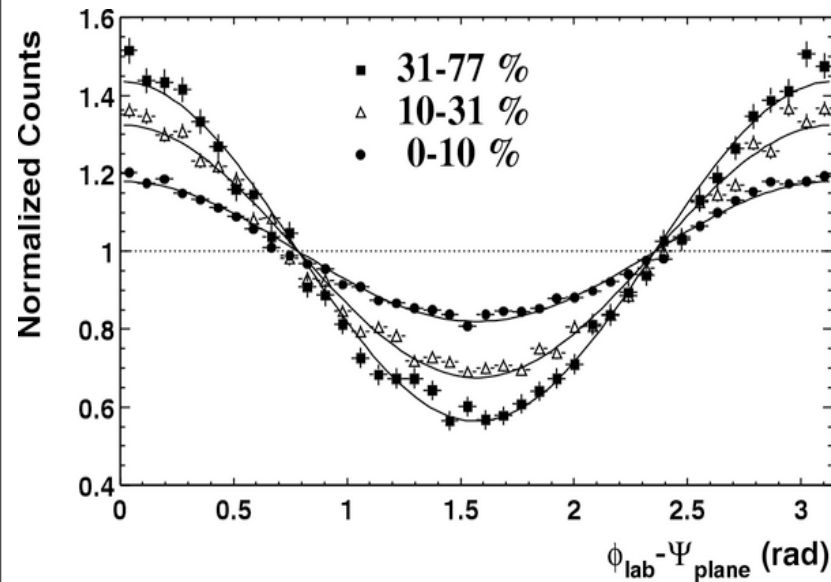


• Very strong elliptic flow → early equilibration

Factor 3:1 peak to valley

Elliptic flow

Distribution of particles with respect to event plane, $\phi-\psi$, $p_t > 2$ GeV; STAR PRL 90 (2003) 032301

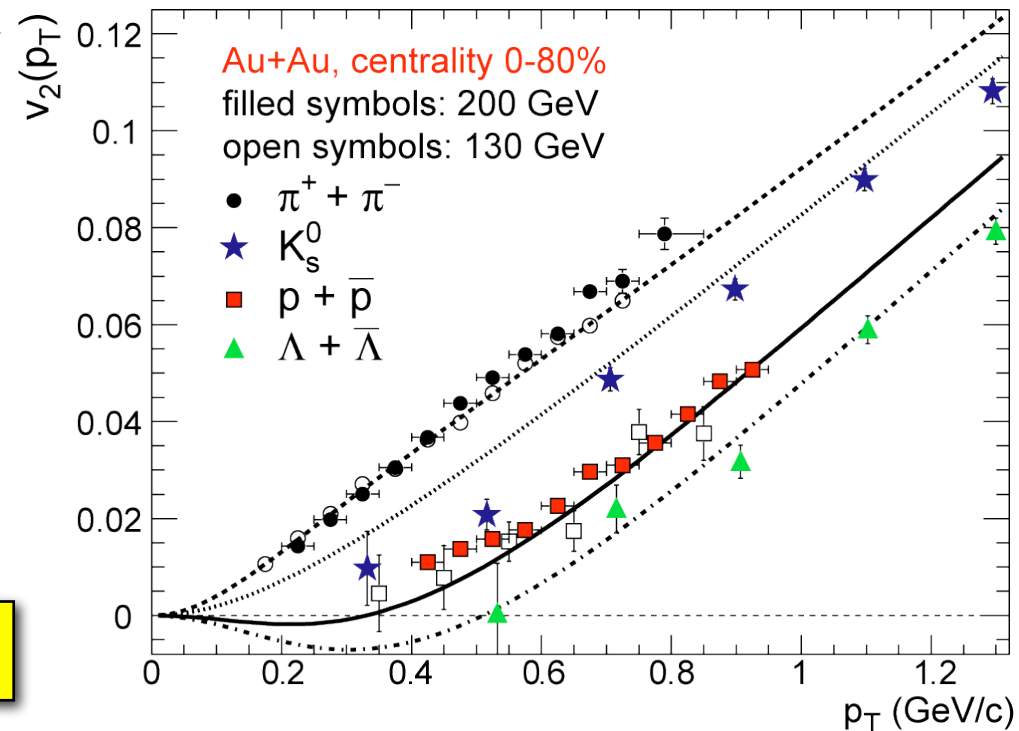


- Pure hydrodynamical models including QGP phase describe elliptic and radial flow for many species

QGP \rightarrow almost perfect fluid

- Very strong elliptic flow \rightarrow early equilibration

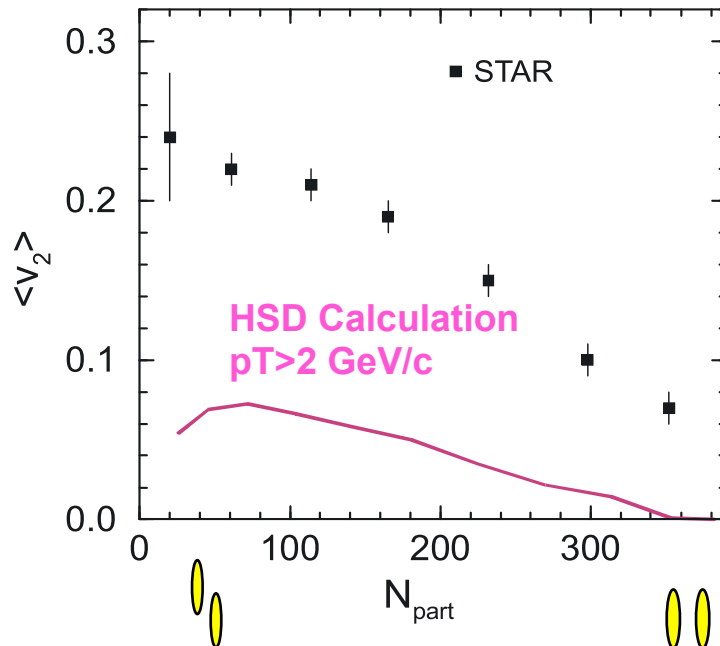
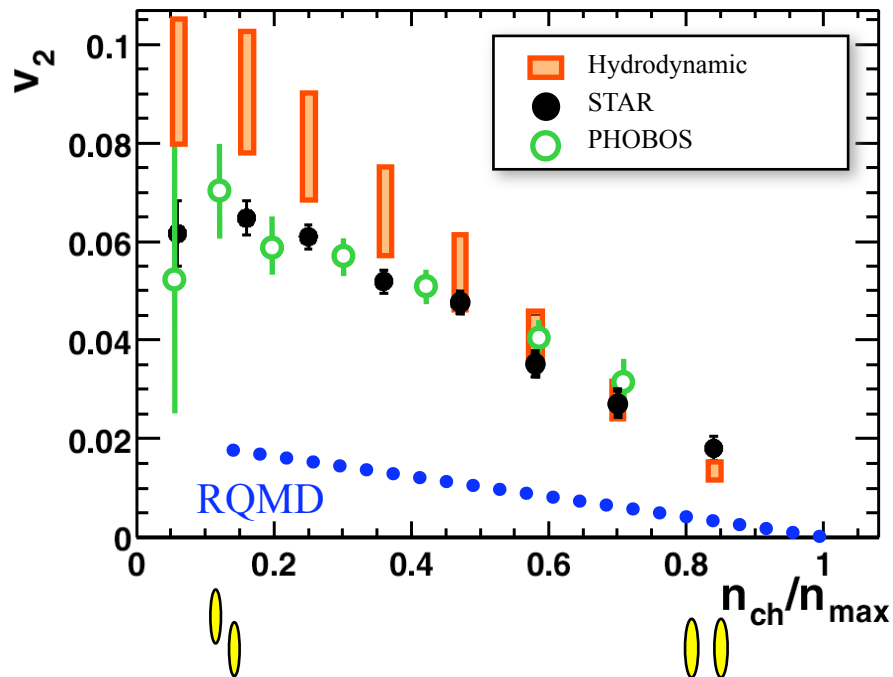
Factor 3:1 peak to valley



Just a gas of hadrons?

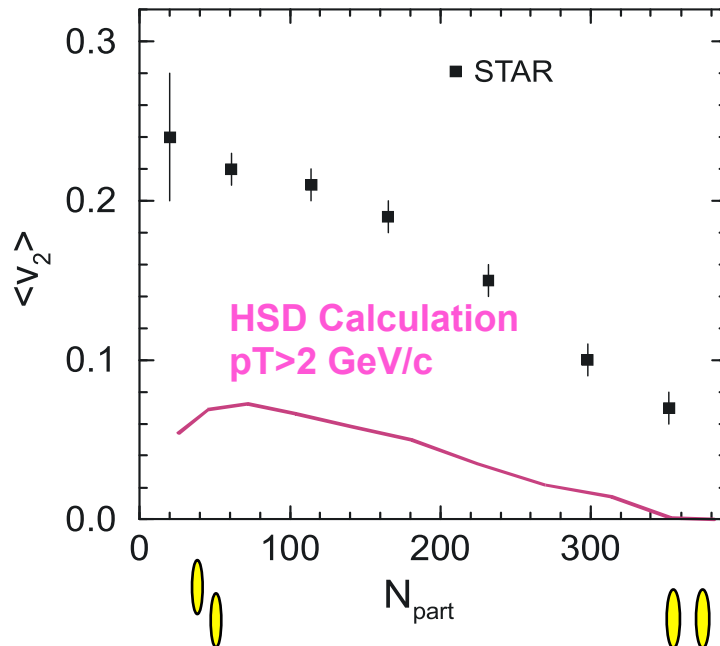
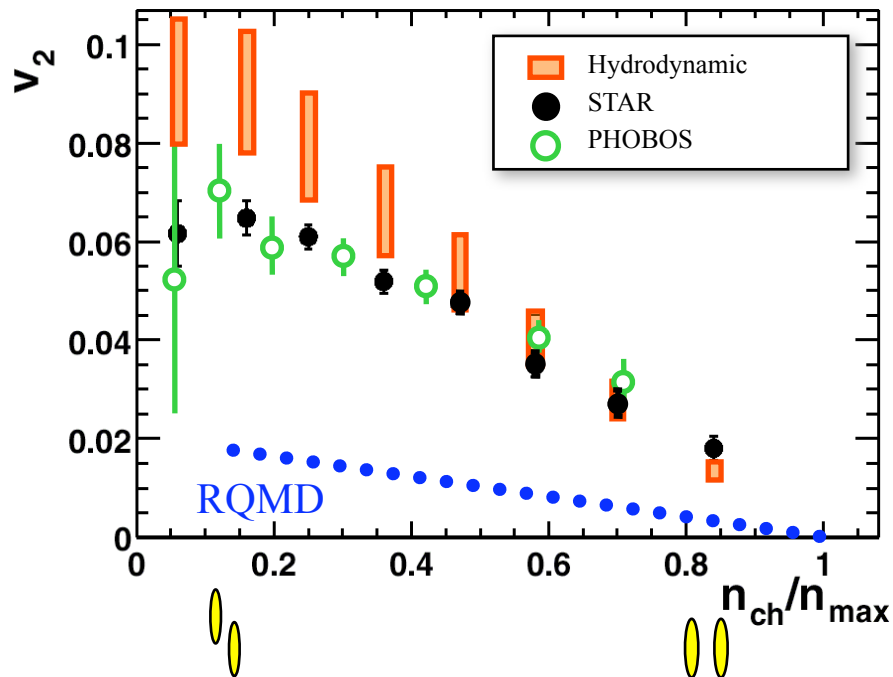
Just a gas of hadrons?

Hadronic transport models (e.g. RQMD, HSD, ...) with hadron formation times ~ 1 fm/c, fail to describe data.



Just a gas of hadrons?

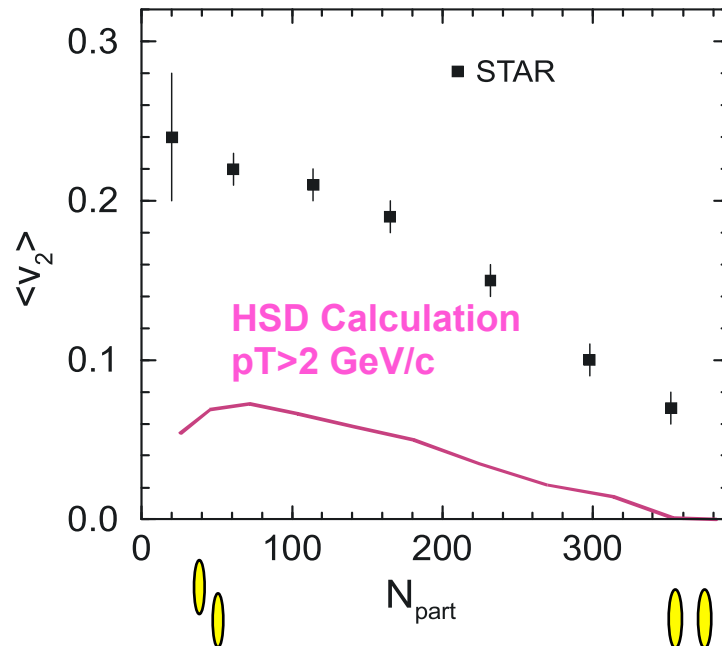
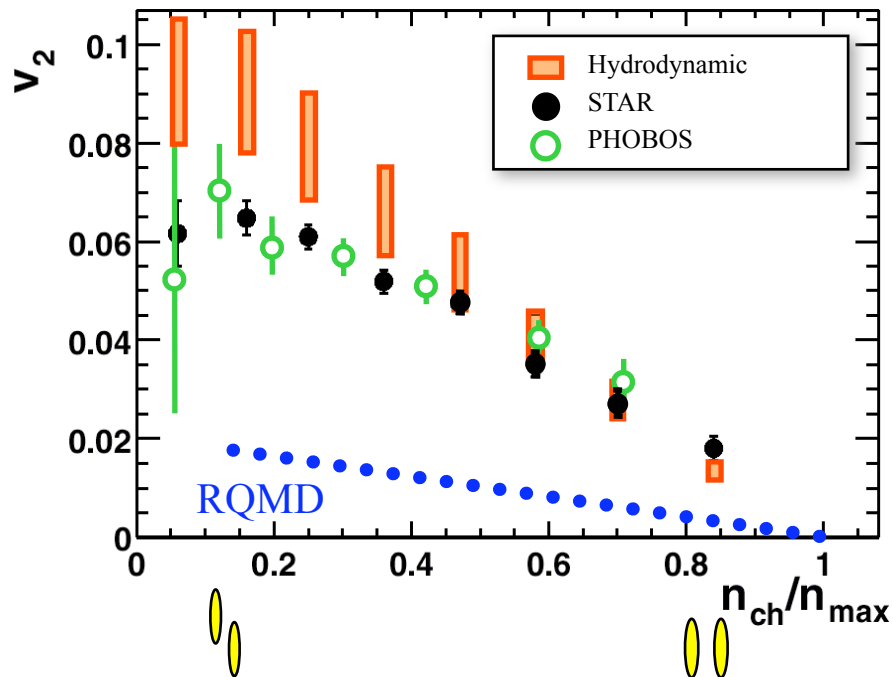
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Clearly the system is not a hadron gas. Not surprising.

Just a gas of hadrons?

Hadronic transport models (e.g. RQMD, HSD, ...) with hadron formation times ~ 1 fm/c, fail to describe data.



Clearly the system is not a hadron gas. Not surprising.

Hydrodynamical calculations: thermalization time $t=0.6$ fm/c

What interactions can lead to equilibration in < 1 fm/c?

The constituents “flow”

- Elliptic flow is additive.
- If partons are flowing the *complicated* observed flow pattern in $v_2(p_T)$ for hadrons

$$\frac{d^2 N}{dp_T d\phi} \propto 1 + 2 v_2(p_T) \cos(2\phi)$$

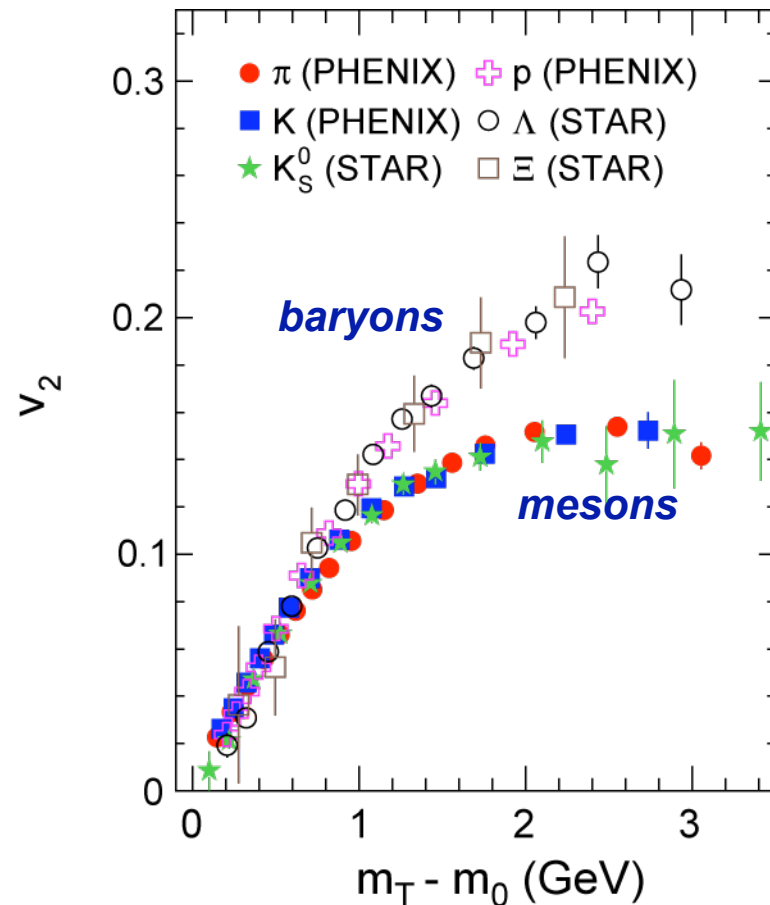
should become *simple* at the quark level

$$p_T \rightarrow p_T / n$$

$$v_2 \rightarrow v_2 / n ,$$

$n = (2, 3)$ for (meson, baryon)

$$m_T = \sqrt{p_T^2 + m_0^2}$$



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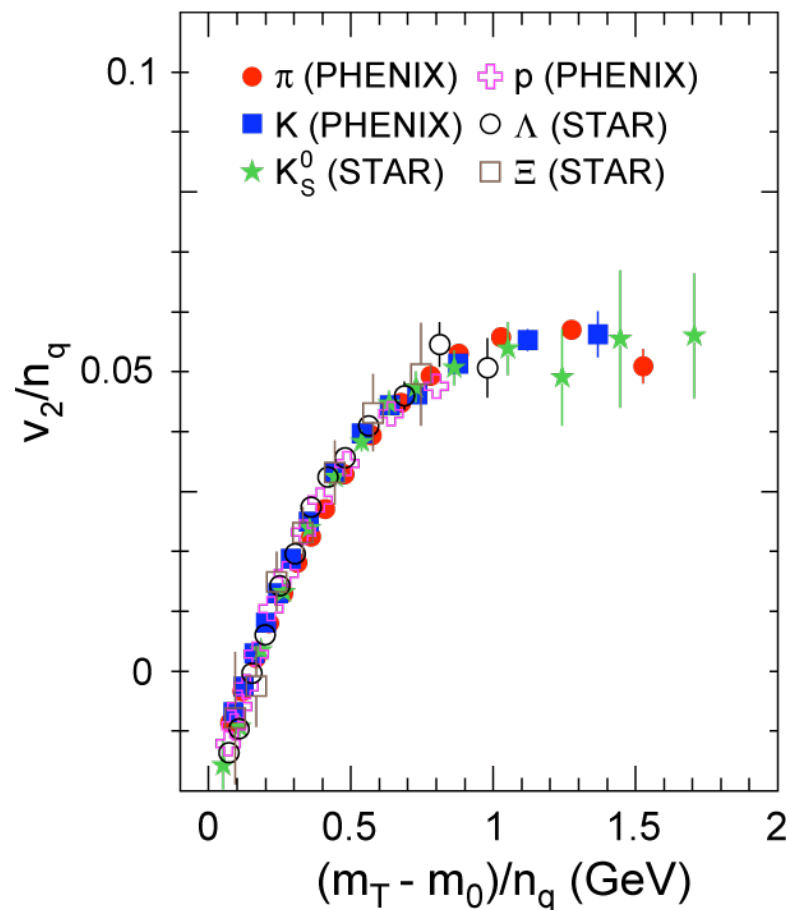
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Works for p , π , K_s^0 , Λ , Ξ ..

$$v_2^s \sim v_2^{u,d} \sim 7\%$$

$$m_T = \sqrt{p_T^2 + m_0^2}$$



Constituents of QGP are partons

Summary of what we learned so far

- Energy density in the collision region is way above that where hadrons can exist
- The initial temperature of collision region is way above that where hadrons can exist
- The medium has quark and gluon degrees of freedom in initial stages

We have created a new state of matter at RHIC
- the QGP

- The QGP is flowing like an almost “perfect” liquid